

Bianchi IX Cosmologies, Spectral Action, and Modular Forms

Matilde Marcolli, Ma148 Spring 2016
Topics in Mathematical Physics

References:

- Wentao Fan, Farzad Fathizadeh, Matilde Marcolli, *Spectral Action for Bianchi type-IX cosmological models*, arXiv:1506.06779, J. High Energy Phys. (2015) 85, 28 pp.
- Wentao Fan, Farzad Fathizadeh, Matilde Marcolli, *Modular forms in the spectral action of Bianchi IX gravitational instantons*, arXiv:1511.05321

Based on Wentao's SURF project with Farzad

$SU(2)$ -Bianchi IX cosmologies (Euclidean, compactified)

- another version of Bianchi IX mixmaster cosmologies, with $SU(2)$ symmetry (Euclidean version)

$$g = w_1 w_2 w_3 dt^2 + \frac{w_2 w_3}{w_1} \sigma_1^2 + \frac{w_1 w_3}{w_2} \sigma_2^2 + \frac{w_1 w_2}{w_3} \sigma_3^2$$

with $w_i = w_i(t)$, or more generally

$$g = F \left(d\mu^2 + \frac{\sigma_1^2}{w_1^2} + \frac{\sigma_2^2}{w_2^2} + \frac{\sigma_3^2}{w_3^2} \right)$$

with a conformal factor $F \sim w_1 w_2 w_3$

- $SU(2)$ -invariant 1-forms $\{\sigma_i\}$ satisfying relations

$$d\sigma_i = \sigma_j \wedge \sigma_k$$

for all cyclic permutations (i, j, k) of $(1, 2, 3)$

- more explicitly ds^2 is

$$\begin{aligned}
 & w_1 w_2 w_3 dt dt + \frac{w_1 w_2 \cos(\eta)}{w_3} d\phi d\psi + \frac{w_1 w_2 \cos(\eta)}{w_3} d\psi d\phi \\
 & + \left(\frac{w_2 w_3 \sin^2(\eta) \cos^2(\psi)}{w_1} + w_1 \left(\frac{w_3 \sin^2(\eta) \sin^2(\psi)}{w_2} + \frac{w_2 \cos^2(\eta)}{w_3} \right) \right) d\phi d\phi \\
 & + \frac{(w_1^2 - w_2^2) w_3 \sin(\eta) \sin(\psi) \cos(\psi)}{w_1 w_2} d\eta d\phi + \frac{(w_1^2 - w_2^2) w_3 \sin(\eta) \sin(\psi) \cos(\psi)}{w_1 w_2} d\phi d\eta \\
 & + \left(\frac{w_2 w_3 \sin^2(\psi)}{w_1} + \frac{w_1 w_3 \cos^2(\psi)}{w_2} \right) d\eta d\eta + \frac{w_1 w_2}{w_3} d\psi d\psi
 \end{aligned}$$

- identifying S^3 with unit quaternions $SU(2)$
- The metrics on S^3

$$\frac{w_2 w_3}{w_1} \sigma_1^2 + \frac{w_1 w_3}{w_2} \sigma_2^2 + \frac{w_1 w_2}{w_3} \sigma_3^2$$

are left-invariants under the action of $SU(2)$ but *not* right-invariant (unlike the round metric on S^3)

Spectral action models of gravity (modified gravity)

- **Spectral triple:** $(\mathcal{A}, \mathcal{H}, D)$
 - 1 unital associative algebra \mathcal{A}
 - 2 represented as bounded operators on a Hilbert space \mathcal{H}
 - 3 Dirac operator: self-adjoint $D^* = D$ with compact resolvent, with bounded commutators $[D, a]$
- prototype: $(C^\infty(M), L^2(M, S), \not{D}_M)$
- extends to non smooth objects (fractals) and noncommutative (NC tori, quantum groups, NC deformations, etc.)

Action functional

- Suppose *finitely summable* $ST = (\mathcal{A}, \mathcal{H}, D)$

$$\zeta_D(s) = \text{Tr}(|D|^{-s}) < \infty, \quad \Re(s) \gg 0$$

- **Spectral action** (Chamseddine–Connes)

$$\mathcal{S}_{ST}(\Lambda) = \text{Tr}(f(D/\Lambda)) = \sum_{\lambda \in \text{Spec}(D)} \text{Mult}(\lambda) f(\lambda/\Lambda)$$

f = smooth approximation to (even) cutoff

Asymptotic expansion (Chamseddine–Connes) for
(almost) commutative geometries:

$$\mathrm{Tr}(f(D/\Lambda)) \sim \sum_{\beta \in \Sigma_{ST}^+} f_\beta \Lambda^\beta \int |D|^{-\beta} + f(0) \zeta_D(0)$$

- Residues

$$\int |D|^{-\beta} = \frac{1}{2} \mathrm{Res}_{s=\beta} \zeta_D(s)$$

- Momenta $f_\beta = \int_0^\infty f(v) v^{\beta-1} dv$
- **Dimension Spectrum** Σ_{ST} poles of zeta functions
 $\zeta_{a,D}(s) = \mathrm{Tr}(a|D|^{-s})$
- positive dimension spectrum $\Sigma_{ST}^+ = \Sigma_{ST} \cap \mathbb{R}_+^*$

Zeta function and heat kernel (manifolds)

- Mellin transform

$$|D|^{-s} = \frac{1}{\Gamma(s/2)} \int_0^\infty e^{-tD^2} t^{\frac{s}{2}-1} dt$$

- heat kernel expansion

$$\mathrm{Tr}(e^{-tD^2}) = \sum_{\alpha} t^{\alpha} c_{\alpha} \quad \text{for } t \rightarrow 0$$

- zeta function expansion

$$\zeta_D(s) = \mathrm{Tr}(|D|^{-s}) = \sum_{\alpha} \frac{c_{\alpha}}{\Gamma(s/2)(\alpha + s/2)} + \text{holomorphic}$$

- taking residues

$$\mathrm{Res}_{s=-2\alpha} \zeta_D(s) = \frac{2c_{\alpha}}{\Gamma(-\alpha)}$$

Pseudo-differential Calculus: (manifold case)

to obtain *full* asymptotic expansion of the Spectral Action

- Dirac operator D and pseudodifferential **symbol** of D^2

$$\sigma(D^2)(x, \xi) = p_2(x, \xi) + p_1(x, \xi) + p_0(x, \xi)$$

each p_k homogeneous of order k in ξ

- **Cauchy integral formula**

$$e^{-tD^2} = \frac{1}{2\pi i} \int_{\gamma} e^{-t\lambda} (D^2 - \lambda)^{-1} d\lambda$$

- **Seeley de-Witt coefficients** ($m = \dim M$)

$$\mathrm{Tr}(e^{-tD^2}) \sim_{t \rightarrow 0^+} t^{-m/2} \sum_{n=0}^{\infty} a_{2n}(D^2) t^n$$

Parametrix Method

- D^2 order 2 elliptic differential operator: exists a parametrix R_λ with

$$\sigma(R_\lambda) \sim \sum_{j=0}^{\infty} r_j(x, \xi, \lambda)$$

- $r_j(x, \xi, \lambda)$ pseudodifferential symbol order $-2 - j$

$$r_j(x, t\xi, t^2\lambda) = t^{-2-j} r_j(x, \xi, \lambda)$$

- R_λ approximates $(D^2 - \lambda)^{-1}$ with $\sigma((D^2 - \lambda)R_\lambda) \sim 1$
- **recursive equation:**

$$\sigma((D^2 - \lambda)R_\lambda) \sim ((p_2(x, \xi) - \lambda) + p_1(x, \xi) + p_0(x, \xi)) \circ \left(\sum_{j=0}^{\infty} r_j(x, \xi, \lambda) \right) \sim 1$$

- **solution** for R_λ constructed recursively:

$$r_0(x, \xi, \lambda) = (p_2(x, \xi) - \lambda)^{-1}$$

$$r_n(x, \xi, \lambda) = - \sum \frac{1}{\alpha!} \partial_\xi^\alpha r_j(x, \xi, \lambda) D_x^\alpha p_k(x, \xi) r_0(x, \xi, \lambda),$$

summation over all $\alpha \in \mathbb{Z}_{\geq 0}^4, j \in \{0, 1, \dots, n-1\}, k \in \{0, 1, 2\}$,
with $|\alpha| + j + 2 - k = n$

Seeley-deWitt coefficients and Parametrix Method

$$a_{2n}(x, D^2) = \frac{(2\pi)^{-m}}{2\pi i} \int \int_\gamma e^{-\lambda} \text{tr}(r_{2n}(x, \xi, \lambda)) d\lambda d^m \xi$$

- odd j coefficients vanish: $r_j(x, \xi, \lambda)$ odd function of ξ

Dirac operator

- orthonormal coframe $\{\theta^a\}$

$$D = \sum_a \theta^a \nabla_{\theta^a}^S$$

- spin connection ∇^S with matrix of 1-forms $\omega = (\omega_b^a)$ with

$$\nabla \theta^a = \sum_b \omega_b^a \otimes \theta^b$$

- metric-compatibility and torsion-freeness (Levi-Civita connection)

$$\omega_b^a = -\omega_a^b, \quad d\theta^a = \sum_b \omega_b^a \wedge \theta^b$$

- Dirac operator

$$D = \sum_{a,\mu} \gamma^a dx^\mu (\theta_a) \frac{\partial}{\partial x^\mu} + \frac{1}{4} \sum_{a,b,c} \gamma^c \omega_{ac}^b \gamma^a \gamma^b$$

with $\omega_a^b = \sum_c \omega_{ac}^b \theta^c$

- matrices γ^a Clifford action of θ^a on spin bundle:

$$(\gamma^a)^2 = -I$$

$$\gamma^a \gamma^b + \gamma^b \gamma^a = 0 \text{ for } a \neq b$$

Dirac operator on Bianchi IX metrics

- local coordinates $(x^\mu) = (t, \eta, \phi, \psi)$ with \mathbb{S}^3 parametrized by

$$(\eta, \phi, \psi) \mapsto \left(\cos(\eta/2)e^{i(\phi+\psi)/2}, \sin(\eta/2)e^{i(\phi-\psi)/2} \right)$$

with $0 \leq \eta \leq \pi, 0 \leq \phi < 2\pi, 0 \leq \psi < 4\pi$.

- orthonormal frame

$$\theta^0 = \sqrt{w_1 w_2 w_3} dt,$$

$$\theta^1 = \sin(\eta) \cos(\psi) \sqrt{\frac{w_2 w_3}{w_1}} d\phi - \sin(\psi) \sqrt{\frac{w_2 w_3}{w_1}} d\eta,$$

$$\theta^2 = \sin(\eta) \sin(\psi) \sqrt{\frac{w_1 w_3}{w_2}} d\phi + \cos(\psi) \sqrt{\frac{w_1 w_3}{w_2}} d\eta,$$

$$\theta^3 = \cos(\eta) \sqrt{\frac{w_1 w_2}{w_3}} d\phi + \sqrt{\frac{w_1 w_2}{w_3}} d\psi.$$

- non-vanishing ω_{ac}^b

$$\omega_{11}^0 = -\frac{w_2 (w_1 w_3' - w_3 w_1') + w_1 w_3 w_2'}{2(w_1 w_2 w_3)^{3/2}}, \quad \omega_{22}^0 = -\frac{w_2 (w_3 w_1' + w_1 w_3') - w_1 w_3 w_2'}{2(w_1 w_2 w_3)^{3/2}},$$

$$\omega_{33}^0 = -\frac{w_2 (w_3 w_1' - w_1 w_3') + w_1 w_3 w_2'}{2(w_1 w_2 w_3)^{3/2}}, \quad \omega_{23}^1 = -\frac{w_1^2 w_2^2 - w_3^2 (w_1^2 + w_2^2)}{2(w_1 w_2 w_3)^{3/2}},$$

$$\omega_{32}^1 = -\frac{w_1^2 (w_2^2 - w_3^2) + w_2^2 w_3^2}{2(w_1 w_2 w_3)^{3/2}}, \quad \omega_{31}^2 = -\frac{w_2^2 w_3^2 - w_1^2 (w_2^2 + w_3^2)}{2(w_1 w_2 w_3)^{3/2}}.$$

pseudo-differential symbol of Dirac

$$\begin{aligned}
 \sigma(D)(x, \xi) &= \sum_{a, \mu} i \gamma^a e_a^\mu \xi_{\mu+1} + \frac{1}{4\sqrt{w_1 w_2 w_3}} \left(\frac{w'_1}{w_1} + \frac{w'_2}{w_2} + \frac{w'_3}{w_3} \right) \gamma^1 \\
 &\quad - \frac{\sqrt{w_1 w_2 w_3}}{4} \left(\frac{1}{w_1^2} + \frac{1}{w_2^2} + \frac{1}{w_3^2} \right) \gamma^2 \gamma^3 \gamma^4 \\
 &= - \frac{i \gamma^2 \sqrt{w_1} (\csc(\eta) \cos(\psi) (\xi_4 \cos(\eta) - \xi_3) + \xi_2 \sin(\psi))}{\sqrt{w_2} \sqrt{w_3}} \\
 &\quad + \frac{i \gamma^3 \sqrt{w_2} (\sin(\psi) (\xi_3 \csc(\eta) - \xi_4 \cot(\eta)) + \xi_2 \cos(\psi))}{\sqrt{w_1} \sqrt{w_3}} \\
 &\quad + \frac{i \gamma^1 \xi_1}{\sqrt{w_1} \sqrt{w_2} \sqrt{w_3}} + \frac{i \gamma^4 \xi_4 \sqrt{w_3}}{\sqrt{w_1} \sqrt{w_2}} \\
 &\quad + \frac{1}{4\sqrt{w_1 w_2 w_3}} \left(\frac{w'_1}{w_1} + \frac{w'_2}{w_2} + \frac{w'_3}{w_3} \right) \gamma^1 \\
 &\quad - \frac{\sqrt{w_1 w_2 w_3}}{4} \left(\frac{1}{w_1^2} + \frac{1}{w_2^2} + \frac{1}{w_3^2} \right) \gamma^2 \gamma^3 \gamma^4
 \end{aligned}$$

- with non-vanishing e_a^μ :

$$e_0^0 = \frac{1}{\sqrt{w_1 w_2 w_3}},$$

$$e_2^1 = \frac{\sqrt{w_2} \cos(\psi)}{\sqrt{w_1 w_3}},$$

$$e_2^2 = \frac{\sqrt{w_2} \csc(\eta) \sin(\psi)}{\sqrt{w_1 w_3}},$$

$$e_2^3 = -\frac{\sqrt{w_2} \cot(\eta) \sin(\psi)}{\sqrt{w_1 w_3}},$$

$$e_1^1 = -\frac{\sqrt{w_1} \sin(\psi)}{\sqrt{w_2 w_3}},$$

$$e_1^2 = \frac{\sqrt{w_1} \csc(\eta) \cos(\psi)}{\sqrt{w_2 w_3}},$$

$$e_1^3 = -\frac{\sqrt{w_1} \cot(\eta) \cos(\psi)}{\sqrt{w_2 w_3}},$$

$$e_3^3 = \frac{\sqrt{w_3}}{\sqrt{w_1 w_2}}$$

- get from the symbol the **homogeneous components** $p_k(x, \xi)$ with

$$\sigma(D^2)(x, \xi) = p_2(x, \xi) + p_1(x, \xi) + p_0(x, \xi)$$

- Example: for $p_0(x, \xi)$ get

$$\begin{aligned} & \left(-\frac{w'_1}{8w_1w_2^2} - \frac{w'_1}{8w_1w_3^2} + \frac{3w'_1}{8w_1^3} - \frac{w'_2}{8w_1^2w_2} - \frac{w'_3}{8w_1^2w_3} - \frac{w'_2}{8w_2w_3^2} \right. \\ & \quad \left. + \frac{3w'_2}{8w_2^3} - \frac{w'_3}{8w_2^2w_3} + \frac{3w'_3}{8w_3^3} \right) \gamma^1 \gamma^2 \gamma^3 \gamma^4 + \\ & \left(-\frac{w''_1}{4w_1^2w_2w_3} + \frac{w'_1w'_2}{8w_1^2w_2^2w_3} + \frac{w'_1w'_3}{8w_1^2w_2w_3^2} + \frac{5w_1'^2}{16w_1^3w_2w_3} - \frac{w''_2}{4w_1w_2^2w_3} \right. \\ & \quad + \frac{w'_2w'_3}{8w_1w_2^2w_3^2} + \frac{5w_2'^2}{16w_1w_2^3w_3} - \frac{w''_3}{4w_1w_2w_3^2} + \frac{5w_3'^2}{16w_1w_2w_3^3} + \frac{w_2w_3}{16w_1^3} \\ & \quad \left. + \frac{w_3}{8w_1w_2} + \frac{w_1w_3}{16w_2^3} + \frac{w_2}{8w_1w_3} + \frac{w_1}{8w_2w_3} + \frac{w_1w_2}{16w_3^3} \right) I. \end{aligned}$$

- also manageable expression for $p_2(x, \xi)$, longer one for $p_1(x, \xi)$

Applying Parametrix Method to this Dirac operator

$$a_{2n}(x, D^2) = \frac{(2\pi)^{-m}}{2\pi i} \int \int_{\gamma} e^{-\lambda} \operatorname{tr}(r_{2n}(x, \xi, \lambda)) d\lambda d^m \xi$$

- Find a_0 , a_2 , a_4 explicitly

$$a_0(D^2) = 4w_1 w_2 w_3$$

$$a_2(D^2) = -\frac{w_1^2}{3} - \frac{w_2^2}{3} - \frac{w_3^2}{3} + \frac{w_1^2 w_2^2}{6w_3^2} + \frac{w_1^2 w_3^2}{6w_2^2} + \frac{w_2^2 w_3^2}{6w_1^2} - \frac{(w_1')^2}{6w_1^2} - \frac{(w_2')^2}{6w_2^2} - \frac{(w_3')^2}{6w_3^2} - \frac{w_1' w_2'}{3w_1 w_2} - \frac{w_1' w_3'}{3w_1 w_3} - \frac{w_2' w_3'}{3w_2 w_3} + \frac{w_1''}{3w_1} + \frac{w_2''}{3w_2} + \frac{w_3''}{3w_3}.$$

and a much longer and more complicated expression for $a_4(D^2)$

Observation: all coefficients in these expressions (also for a_4) are **rational numbers** ... what about other terms in expansion?

A different method: **Wodzicki residue**

- **Wodzicki residue**: unique trace functional on algebra of pseudodifferential operators on smooth sections of vector bundle over smooth manifold
- classical pseudodifferential operator P_σ of order $d \in \mathbb{Z}$ local symbol

$$\sigma(x, \xi) \sim \sum_{j=0}^{\infty} \sigma_{d-j}(x, \xi) \quad (\xi \rightarrow \infty),$$

σ_{d-j} positively homogeneous order $d - j$ in ξ

- **Residue**:

$$\text{Res}(P_\sigma) = \int_{S^*M} \text{Tr}(\sigma_{-m}(x, \xi)) d^{m-1}\xi d^m x,$$

$S^*M = \{(x, \xi) \in T^*M; \|\xi\|_g = 1\}$ cosphere bundle

- **spectral formulation** of residue: pseudodifferential operator P_σ , Laplacian Δ

$$P_\sigma \mapsto \text{Res}_{s=0} \text{Tr}(P_\sigma \Delta^{-s})$$

same up to a constant $c_m = 2^{m+1} \pi^m$

- **Mellin transform** (for simplicity $\text{Ker}(\Delta) = 0$):

$$\text{Tr}(\Delta^{-s}) = \frac{1}{\Gamma(s)} \int_0^\infty \text{Tr}(e^{-t\Delta}) t^s \frac{dt}{t}$$

- **heat kernel expansion**

$$\text{Tr}(e^{-t\Delta}) = t^{-m/2} \sum_{n=0}^N a_{2n} t^n + O(t^{-m/2+N+1})$$

- find for any non-negative integer $n \leq m/2 - 1$:

$$\text{Res}_{s=m/2-n} \text{Tr}(\Delta^{-s}) = \frac{a_{2n}(\Delta)}{\Gamma(m/2 - n)},$$

- in particular

$$\text{Res}_{s=1} \text{Tr}(\Delta^{-s}) = a_{m-2}(\Delta)$$

- in terms of **Wodicki residue**:

$$a_{m-2}(\Delta) = \frac{1}{c_m} \text{Res}(\Delta^{-1}) = \frac{1}{2^{m+1} \pi^m} \text{Res}(\Delta^{-1})$$

applied to $\Delta = D^2$

- coefficient $a_2(D^2)$

$$a_2(D^2) = \frac{1}{c_4} \text{Res}(D^{-2}) = \frac{1}{32\pi^4} \int_{S^*M} \text{Tr}(\sigma_{-4}(D^{-2})) d^3\xi d^4x$$

- for other coefficients, introduce an **auxiliary product space** for correct counting of dimensions: use flat r -dimensional torus $\mathbb{T}^r = (\mathbb{R}/\mathbb{Z})^r$

$$\Delta = D^2 \otimes 1 + 1 \otimes \Delta_{\mathbb{T}^r},$$

$\Delta_{\mathbb{T}^r}$ flat Laplacian on \mathbb{T}^r

$$a_{2+r}(D^2) = \frac{1}{2^r \pi^{4+r/2}} \text{Res}(\Delta^{-1})$$

because Künneth formula gives

$$a_{2+r}((x, x'), \Delta) = a_{2+r}(x, D^2) a_0(x', \mathbb{T}^r) = 2^{-r} \pi^{-r/2} a_{2+r}(x, D^2)$$

with volume term only non-zero heat coefficient for flat metric

- obtain for **all coefficients**

$$a_{2+r}(D^2) = \frac{1}{2^5 \pi^{4+r/2}} \int \text{Tr}(\sigma_{-4-r}(\Delta^{-1})) d^{3+r} \xi d^4 x.$$

- writing $\sigma(\Delta^{-1}) \sim \sum_{j=-2}^{-\infty} \sigma_j(x, \xi)$ inductively

$$\sigma_{-2}(x, \xi) = p'_2(x, \xi)^{-1},$$

$$\sigma_{-2-n}(x, \xi) = - \sum \frac{1}{\alpha!} \partial_\xi^\alpha \sigma_j(x, \xi) D_x^\alpha p_k(x, \xi) \sigma_{-2}(x, \xi) \quad (n > 0),$$

summation over all multi-indices non-negative integers α ,
 $-2 - n < j \leq -2, 0 \leq k \leq 2$, with $|\alpha| - j - k = n$

- setting $\zeta_{\mu+1} = \sum_{\nu} e_{\mu}^{\nu} \xi_{\nu+1}$ find inductively for $n \geq 2$

$$\sigma_{-2-n}(x, \xi)|_{S^*(M \times \mathbb{T}^{n-2})} = \sigma_{-2-n}(x, \xi(\zeta))|_{\zeta \in \mathbb{S}^{n+1}} = (w_1 w_2 w_3)^{-\frac{3}{2}n} P_n(\zeta)$$

polynomials $P_n(\zeta)$ coefficients functions of w_i and derivatives

- these explicitly give

$$a_{2n}(D^2) = (w_1 w_2 w_3)^{1-3n} Q_{2n} \left(w_1, w_2, w_3, w'_1, w'_2, w'_3, \dots, w_1^{(2n)}, w_2^{(2n)}, w_3^{(2n)} \right)$$

with Q_{2n} polynomials with **rational coefficients**

$$Q_{2n} = \frac{1}{2\pi^{n+1}} \int_{\mathbb{S}^{2n+1}} \text{Tr}(P_{2n}(\zeta)(\Delta^{-1})) d^{2n+1}\zeta$$

Question: is this rationality a sign of an **arithmetic structure** of Bianchi IX gravitational instantons that persists in the Spectral Action?

Blanchi IX gravitational instantons and Painlevé VI

- Euclidean Bianchi IX metrics with $SU(2)$ -symmetry that are
 - self-dual (Weyl curvature tensor W self-dual)
 - Einstein metrics (Ricci tensor proportional to the metric)
- Self-dual equations for a Riemannian 4-manifold are PDEs; with $SU(2)$ -symmetry reduce to ODEs
- This ODE is a Painlevé VI equation with

$$(\alpha, \beta, \gamma, \delta) = \left(\frac{1}{8}, -\frac{1}{8}, \frac{1}{8}, \frac{3}{8}\right)$$

- N.J. Hitchin. *Twistor spaces, Einstein metrics and isomonodromic deformations*, J.Diff.Geom., Vol. 42, No. 1 (1995), 30–112.
- K.P. Tod. *Self-dual Einstein metrics from the Painlevé VI equation*, Phys. Lett. A 190 (1994), 221–224.
- S. Okumura. *The self-dual Einstein–Weyl metric and classical solutions of Painlevé VI*, Lett. in Math. Phys., 46 (1998), 219–232.
- M.V. Babich, D.A. Korotkin, *Self-dual $SU(2)$ -Invariant Einstein Metrics and Modular Dependence of Theta-Functions*. Lett. Math. Phys. 46 (1998), 323–337

Painlevé VI equations

- *Painlevé transcendents*: solutions of nonlinear second-order ODEs in the plane with *Painlevé property* (the only movable singularities are poles) not solvable in terms of elementary functions; classification in types
- *Painlevé VI*: 4-parameter family $(\alpha, \beta, \gamma, \delta)$

$$\begin{aligned} \frac{d^2 X}{dt^2} = & \frac{1}{2} \left(\frac{1}{X} + \frac{1}{X-1} + \frac{1}{X-t} \right) \left(\frac{dX}{dt} \right)^2 \\ & - \left(\frac{1}{t} + \frac{1}{t-1} + \frac{1}{X-t} \right) \frac{dX}{dt} + \\ & + \frac{X(X-1)(X-t)}{t^2(t-1)^2} \left(\alpha + \beta \frac{t}{X^2} + \gamma \frac{t-1}{(X-1)^2} + \delta \frac{t(t-1)}{(X-t)^2} \right). \end{aligned}$$

Painlevé VI and elliptic curves

- Painlevé VI rewritten as (Fuchs)

$$t(1-t) \left[t(1-t) \frac{d^2}{dt^2} + (1-2t) \frac{d}{dt} - \frac{1}{4} \right] \int_{\infty}^{(X,Y)} \frac{dx}{\sqrt{x(x-1)(x-t)}} = \\ = \alpha Y + \beta \frac{tY}{X^2} + \gamma \frac{(t-1)Y}{(X-1)^2} + \left(\delta - \frac{1}{2} \right) \frac{t(t-1)Y}{(X-t)^2}$$

where $(X, Y) := (X(t), Y(t))$ is a section

(local and/or multivalued) $P := (X(t), Y(t))$

of the generic elliptic curve $E = E(t) : Y^2 = X(X-1)(X-t)$

- left-hand-side $\mu(P)$ satisfies $\mu(P+Q) = \mu(P) + \mu(Q)$ for $P+Q$ addition on the elliptic curve E (in particular $\mu(Q) = 0$ for points of finite order)

- analytic description of the elliptic curve $E_\tau = \mathbb{C}/\Lambda$ with $\Lambda = \mathbb{Z} + \tau\mathbb{Z}$, with $\tau \in \mathbb{H}$
- then Painlevé VI rewritten as (Manin)

$$\frac{d^2 z}{d\tau^2} = \frac{1}{(2\pi i)^2} \sum_{j=0}^3 \alpha_j \wp_z\left(z + \frac{T_j}{2}, \tau\right)$$

with $(\alpha_0, \dots, \alpha_3) := (\alpha, -\beta, \gamma, \frac{1}{2} - \delta)$ and $(T_0, T_1, T_2, T_3) := (0, 1, \tau, 1 + \tau)$, and

$$\wp(z, \tau) := \frac{1}{z^2} + \sum_{(m,n) \neq (0,0)} \left(\frac{1}{(z - m\tau - n)^2} - \frac{1}{(m\tau + n)^2} \right)$$

- also have, for $e_i(\tau) = \wp(\frac{T_i}{2}, \tau)$

$$\wp_z(z, \tau)^2 = 4(\wp(z, \tau) - e_1(\tau))(\wp(z, \tau) - e_2(\tau))(\wp(z, \tau) - e_3(\tau))$$

so $e_1 + e_2 + e_3 = 0$

- a multivalued solution $z = z(\tau)$ defines a multi-section of the family, which is a covering of \mathbb{H}
- is know ramification and monodromy can study behavior over geodesics in \mathbb{H}

- Yu.I. Manin, *Sixth Painlevé equation, universal elliptic curve, and mirror of \mathbb{P}^2* , in “Geometry of Differential Equations”, Amer. Math. Soc. Transl. (2) Vol. 186 (1998) 131–151

Theta characteristics

- explicit parameterization of solutions for coefficients W_i of the Bianchi IX gravitational instantons (from solutions of Painlevé VI)
- **theta-characteristics** with parameters (p, q) :

$$\vartheta[p, q](z, i\mu) := \sum_{m \in \mathbb{Z}} \exp(-\pi(m+p)^2\mu + 2\pi i(m+p)(z+q))$$

- theta-characteristics and theta functions with vanishing characteristics

$$\vartheta[p, q](z, i\mu) = \exp(-\pi p^2\mu + 2\pi ipq) \cdot \vartheta[0, 0](z + pi\mu + q, i\mu)$$

Gravitational instantons and theta characteristics

- use notation $\vartheta[p, q] := \vartheta[p, q](0, i\mu)$, and

$$\vartheta_2 := \vartheta[1/2, 0], \quad \vartheta_3 := \vartheta[0, 0], \quad \vartheta_4 := \vartheta[0, 1/2]$$

- self-dual metrics

$$g = F \left(d\mu^2 + \frac{\sigma_1^2}{w_1^2} + \frac{\sigma_2^2}{w_2^2} + \frac{\sigma_3^2}{w_3^2} \right)$$

with

$$w_1 = -\frac{i}{2} \vartheta_3 \vartheta_4 \frac{\frac{\partial}{\partial q} \vartheta[p, q + \frac{1}{2}]}{e^{\pi i p} \vartheta[p, q]}, \quad w_2 = \frac{i}{2} \vartheta_2 \vartheta_4 \frac{\frac{\partial}{\partial q} \vartheta[p + \frac{1}{2}, q + \frac{1}{2}]}{e^{\pi i p} \vartheta[p, q]},$$

$$w_3 = -\frac{1}{2} \vartheta_2 \vartheta_3 \frac{\frac{\partial}{\partial q} \vartheta[p + \frac{1}{2}, q]}{\vartheta[p, q]},$$

- with non-zero cosmological constant Λ :

$$F = \frac{2}{\pi \Lambda} \frac{w_1 w_2 w_3}{\left(\frac{\partial}{\partial q} \log \vartheta[p, q] \right)^2}$$

- these metrics also satisfy **Einstein equation** if either
 - ① $\Lambda < 0$ with $p \in \mathbb{R}$ and $q \in \frac{1}{2} + i\mathbb{R}$
 - ② $\Lambda > 0$ with $q \in \mathbb{R}$ and $p \in \frac{1}{2} + i\mathbb{R}$
- also case with **vanishing cosmological constant**:

$$w_1 = \frac{1}{\mu + q_0} + 2 \frac{d}{d\mu} \log \vartheta_2, \quad w_2 = \frac{1}{\mu + q_0} + 2 \frac{d}{d\mu} \log \vartheta_3,$$

$$w_3 = \frac{1}{\mu + q_0} + 2 \frac{d}{d\mu} \log \vartheta_4, \quad F = C(\mu + q_0)^2 w_1 w_2 w_3$$

with $q_0, C \in \mathbb{R}, C > 0$.

- M.V. Babich, D.A. Korotkin, *Self-dual $SU(2)$ -Invariant Einstein Metrics and Modular Dependence of Theta-Functions*. Lett. Math. Phys. 46 (1998), 323–337
- Yuri Manin, Matilde Marcolli, *Symbolic Dynamics, Modular Curves, and Bianchi IX Cosmologies*, arXiv:1504.04005 [gr-qc]

Bianchi IX: time-dependent conformal perturbations

- original **triaxial Bianchi IX**:

$$ds^2 = w_1 w_2 w_3 d\mu^2 + \frac{w_2 w_3}{w_1} \sigma_1^2 + \frac{w_3 w_1}{w_2} \sigma_2^2 + \frac{w_1 w_2}{w_3} \sigma_3^2$$

$w_i = w_i(\mu)$ cosmic time μ

- time-dependent **conformal perturbation**:

$$d\tilde{s}^2 = F ds^2 = F \left(w_1 w_2 w_3 d\mu^2 + \frac{w_2 w_3}{w_1} \sigma_1^2 + \frac{w_3 w_1}{w_2} \sigma_2^2 + \frac{w_1 w_2}{w_3} \sigma_3^2 \right)$$

with $F = F(\mu)$

- effect on **Dirac operator**:

$$\tilde{D} = \frac{1}{\sqrt{F}} D + \frac{3F'}{4F^{\frac{3}{2}} w_1 w_2 w_3} \gamma^0$$

D Dirac operator of unperturbed Bianchi IX

- **spectral action** expansion for \tilde{D} from **heat kernel**

$$\mathrm{Tr} \left(\exp(-t\tilde{D}^2) \right) \sim t^{-2} \sum_{n=0}^{\infty} \tilde{a}_{2n} t^n, \quad t \rightarrow 0^+$$

- **rationality** result for coefficients of the spectral action

$$\tilde{a}_{2n} = \frac{\tilde{Q}_{2n} \left(w_1, w_2, w_3, F, w'_1, w'_2, w'_3, F', \dots, w_1^{(2n)}, w_2^{(2n)}, w_3^{(2n)}, F^{(2n)} \right)}{F^{2n} (w_1 w_2 w_3)^{3n-1}}$$

\tilde{Q}_{2n} polynomial with rational coefficients

- zeroth coefficient: volume form (cosmological term)

$$\tilde{a}_0 = 4F^2 w_1 w_2 w_3$$

- second coefficient \tilde{a}_2 : Einstein-Hilbert action

$$-\frac{F}{3} (w_1^2 + w_2^2 + w_3^2) + \frac{F}{6} \left(\frac{w_1^2 w_2^2 - w_3'^2}{w_3^2} + \frac{w_1^2 w_3^2 - w_2'^2}{w_2^2} + \frac{w_2^2 w_3^2 - w_1'^2}{w_1^2} \right)$$

$$-\frac{F}{3} \left(\frac{w_1' w_2'}{w_1 w_2} + \frac{w_1' w_3'}{w_1 w_3} + \frac{w_2' w_3'}{w_2 w_3} \right) + \frac{F}{3} \left(\frac{w_1''}{w_1} + \frac{w_2''}{w_2} + \frac{w_3''}{w_3} \right) - \frac{F'^2}{2F} + F''$$

- much longer and more complicated explicit formula for \tilde{a}_4
(Weyl conformal gravity and Gauss-Bonnet gravity)

Gravitational Instantons

- now assuming conformally perturbed Bianchi IX is **self-dual Einstein metric** and use parameterization by **theta functions**
- two-parameter family with non-vanishing cosmological constant:

$$w_1[p, q](i\mu) = -\frac{i}{2}\vartheta_3(i\mu)\vartheta_4(i\mu) \frac{\partial_q \vartheta[p, q + \frac{1}{2}](i\mu)}{e^{\pi i p} \vartheta[p, q](i\mu)}$$

$$w_2[p, q](i\mu) = \frac{i}{2}\vartheta_2(i\mu)\vartheta_4(i\mu) \frac{\partial_q \vartheta[p + \frac{1}{2}, q + \frac{1}{2}](i\mu)}{e^{\pi i p} \vartheta[p, q](i\mu)}$$

$$w_3[p, q](i\mu) = -\frac{1}{2}\vartheta_2(i\mu)\vartheta_3(i\mu) \frac{\partial_q \vartheta[p + \frac{1}{2}, q](i\mu)}{\vartheta[p, q](i\mu)}$$

$$F[p, q](i\mu) = \frac{2}{\pi\Lambda} \frac{1}{(\partial_q \ln \vartheta[p, q](i\mu))^2} = \frac{2}{\pi\Lambda} \left(\frac{\vartheta[p, q](i\mu)}{\partial_q \vartheta[p, q](i\mu)} \right)^2$$

- one-parameter family with vanishing cosmological constant:

$$w_1[q_0](i\mu) = \frac{1}{\mu + q_0} + 2\frac{d}{d\mu} \log \vartheta_2(i\mu),$$

$$w_2[q_0](i\mu) = \frac{1}{\mu + q_0} + 2\frac{d}{d\mu} \log \vartheta_3(i\mu),$$

$$w_3[q_0](i\mu) = \frac{1}{\mu + q_0} + 2\frac{d}{d\mu} \log \vartheta_4(i\mu),$$

$$F[q_0](i\mu) = C(\mu + q_0)^2,$$

C arbitrary positive constant

Modularity

- generators of the modular group $\mathrm{PSL}_2(\mathbb{Z})$

$$T_1(\tau) = \tau + 1, \quad S(\tau) = \frac{-1}{\tau}, \quad \tau \in \mathbb{H}$$

- using behavior of theta functions and derivatives under modular transformations (two-parameter family):

$$\begin{aligned} w_1[p, q](i\mu + 1) &= w_1[p, q + p + \frac{1}{2}](i\mu), & w_1^{(n)}[p, q](i\mu + 1) &= w_1^{(n)}[p, q + p + \frac{1}{2}](i\mu), \\ w_2[p, q](i\mu + 1) &= w_3[p, q + p + \frac{1}{2}](i\mu), & w_2^{(n)}[p, q](i\mu + 1) &= w_3^{(n)}[p, q + p + \frac{1}{2}](i\mu), \\ w_3[p, q](i\mu + 1) &= w_2[p, q + p + \frac{1}{2}](i\mu), & w_3^{(n)}[p, q](i\mu + 1) &= w_2^{(n)}[p, q + p + \frac{1}{2}](i\mu). \end{aligned}$$

- for μ with $\Re(\mu) > 0$:

$$w_3[p, q]\left(\frac{i}{\mu}\right) = -\mu^2 w_1[-q, p](i\mu),$$

$$w_3'[p, q]\left(\frac{i}{\mu}\right) = \mu^4 w_1'[-q, p](i\mu) + 2\mu^3 w_1[-q, p](i\mu),$$

$$w_3''[p, q]\left(\frac{i}{\mu}\right) = -\mu^6 w_1''[-q, p](i\mu) - 6\mu^5 w_1'[-q, p](i\mu) - 6\mu^4 w_1[-q, p](i\mu),$$

$$w_3^{(3)}[p, q]\left(\frac{i}{\mu}\right) = \mu^8 w_1^{(3)}[-q, p](i\mu) + 12\mu^7 w_1''[-q, p](i\mu) + 36\mu^6 w_1'[-q, p](i\mu) + 24\mu^5 w_1[-q, p](i\mu),$$

$$w_3^{(4)}[p, q]\left(\frac{i}{\mu}\right) = -\mu^{10} w_1^{(4)}[-q, p](i\mu) - 20\mu^9 w_1^{(3)}[-q, p](i\mu) - 120\mu^8 w_1''[-q, p](i\mu) - 240\mu^7 w_1'[-q, p](i\mu) - 120\mu^6 w_1[-q, p](i\mu).$$

- similar results for w_2 and w_3 under modular generator S

- conformal factor:

$$F[p, q](i\mu + 1) = F[p, q + p + \frac{1}{2}](i\mu),$$

$$F^{(n)}[p, q](i\mu + 1) = F^{(n)}[p, q + p + \frac{1}{2}](i\mu).$$

$$F[p, q]\left(\frac{i}{\mu}\right) = -\mu^{-2}F[-q, p](i\mu),$$

$$F'[p, q]\left(\frac{i}{\mu}\right) = F'[-q, p](i\mu) - 2\mu^{-1}F[-q, p](i\mu),$$

$$F''[p, q]\left(\frac{i}{\mu}\right) = -\mu^2 F''[-q, p](i\mu) + 2\mu F'[-q, p](i\mu) - 2F[-q, p](i\mu),$$

$$F^{(3)}[p, q]\left(\frac{i}{\mu}\right) = \mu^4 F^{(3)}[-q, p](i\mu),$$

$$F^{(4)}[p, q]\left(\frac{i}{\mu}\right) = -\mu^6 F^{(4)}[-q, p](i\mu) - 4\mu^5 F^{(3)}[-q, p](i\mu).$$

- similar results for the case of the one-parameter family with vanishing cosmological constant
- **modularity of spectral action coefficients:**

$$\tilde{a}_0[p, q](i\mu + 1) = \tilde{a}_0[p, q + p + \frac{1}{2}](i\mu)$$

$$\tilde{a}_2[p, q](i\mu + 1) = \tilde{a}_2[p, q + p + \frac{1}{2}](i\mu)$$

$$\tilde{a}_4[p, q](i\mu + 1) = \tilde{a}_4[p, q + p + \frac{1}{2}](i\mu)$$

$$\tilde{a}_0[p, q]\left(\frac{i}{\mu}\right) = -\mu^2 \tilde{a}_0[-q, p](i\mu)$$

$$\tilde{a}_2[p, q]\left(\frac{i}{\mu}\right) = -\mu^2 \tilde{a}_2[-q, p](i\mu)$$

$$\tilde{a}_4[p, q]\left(\frac{i}{\mu}\right) = -\mu^2 \tilde{a}_4[-q, p](i\mu)$$

Modularity of remaining coefficients \tilde{a}_{2n}

- Dirac operators $\tilde{D}^2[p, q]$, $\tilde{D}^2[p, q + p + \frac{1}{2}]$ and $\tilde{D}^2[-q, p]$ are **isospectral**
- heat kernel $K_t[p, q]$ of $\exp(-t\tilde{D}^2[p, q])$ in terms of eigenvalues and eigenspinors \Rightarrow modularity

$$K_t[p, q](i\mu_1 + 1, i\mu_2 + 1) = K_t[p, q + p + \frac{1}{2}](i\mu_1, i\mu_2),$$

$$K_t[p, q]\left(-\frac{1}{i\mu_1}, -\frac{1}{i\mu_2}\right) = (i\mu_2)^2 K_t[-q, p](i\mu_1, i\mu_2).$$

- then modularity of coefficients \tilde{a}_{2n} :

$$\tilde{a}_{2n}[p, q](i\mu + 1) = \tilde{a}_{2n}[p, q + p + \frac{1}{2}](i\mu),$$

$$\tilde{a}_{2n}[p, q]\left(\frac{i}{\mu}\right) = (i\mu)^2 \tilde{a}_{2n}[-q, p](i\mu).$$

Vector valued modular forms

- coefficients satisfy:

$$\tilde{a}_{2n}[p+1, q] = \tilde{a}_{2n}[p, q+1] = \tilde{a}_{2n}[p, q],$$

- $\mathrm{PSL}_2(\mathbb{Z})$ action on $(p, q) \in \mathbb{R}/\mathbb{Z}^2$:

$$\begin{aligned}\tilde{S}(p, q) &= (-q, p) \\ \tilde{T}_1(p, q) &= (p, q + p + \frac{1}{2})\end{aligned}$$

finite orbits $\mathcal{O}_{(p,q)}$ on rationals

- $\tilde{a}_{2n}[p', q'](i\mu)$, with $(p', q') \in \mathcal{O}_{(p,q)}$, **vector-valued modular form** of weight 2 for the modular group $\mathrm{PSL}_2(\mathbb{Z})$

- summing over orbits:

$$\tilde{a}_{2n}(i\mu; \mathcal{O}_{(p,q)}) = \sum_{(p',q') \in \mathcal{O}_{(p,q)}} \tilde{a}_{2n}[p', q'](i\mu)$$

is an ordinary **modular form** of weight 2 for $\mathrm{PSL}_2(\mathbb{Z})$

- **Question:** which modular form is it?
- analyze zeros and poles structure to find out
 - **Example:** for all n , modular form $\tilde{a}_{2n}(i\mu; \mathcal{O}_{(0, \frac{1}{3})})$ in one-dimensional space spanned by

$$\frac{G_{14}(i\mu)}{\Delta(i\mu)},$$

with Δ modular discriminant (cusp form weight 12) and G_{14} is Eisenstein series weight 14

- **Example:** for all n , modular form $\tilde{a}_{2n}(i\mu; \mathcal{O}_{(\frac{1}{6}, \frac{5}{6})})$ in one-dimensional space spanned by

$$\frac{\Delta(i\mu)G_6(i\mu)}{G_4(i\mu)^4}$$