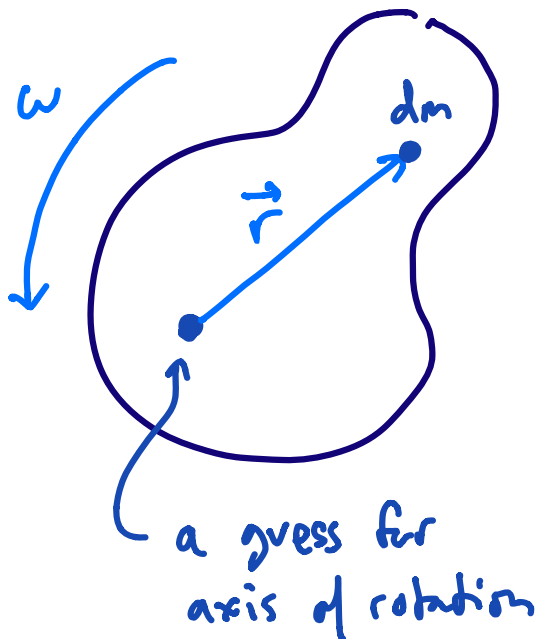


Lecture 11: Rotational Dynamics II

Free Rotation

Consider an object spinning off in empty space.



$$\text{Centripetal force on } dm = d\vec{F} = -dm \omega^2 \vec{r}$$

$$\text{no external force} \Rightarrow \vec{0} = \vec{F} = \int d\vec{F} = -\omega^2 \int dm \vec{r}$$

$$\int dm \vec{r} = 0 \text{ implies that } \boxed{\text{Pivot axis} = \text{CoM axis}}$$

bottom line: objects in empty space rotate about their center of mass

Physical Pendulum



gravity acts to decrease θ

$$\tau = -mgr \sin \theta = I \ddot{\theta} = \frac{d^2 \theta}{dt^2}$$

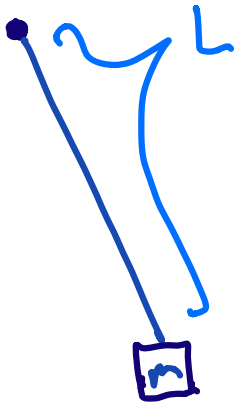
for small $\theta \ll 1$,

$$\ddot{\theta} = -\frac{mgr}{I} \theta$$

we will see later, this defines oscillatory motion w/ the period

$$T = 2\pi \sqrt{\frac{I}{mgr}}$$

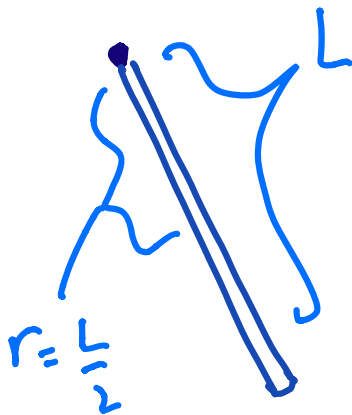
Compare a simple versus stick pendulum.



$$I_{\text{simp}} = mL^2 \text{ and } r = L$$



$$T_{\text{simp}} = 2\pi \sqrt{\frac{mL^2}{mgL}} = 2\pi \sqrt{\frac{L}{g}}$$

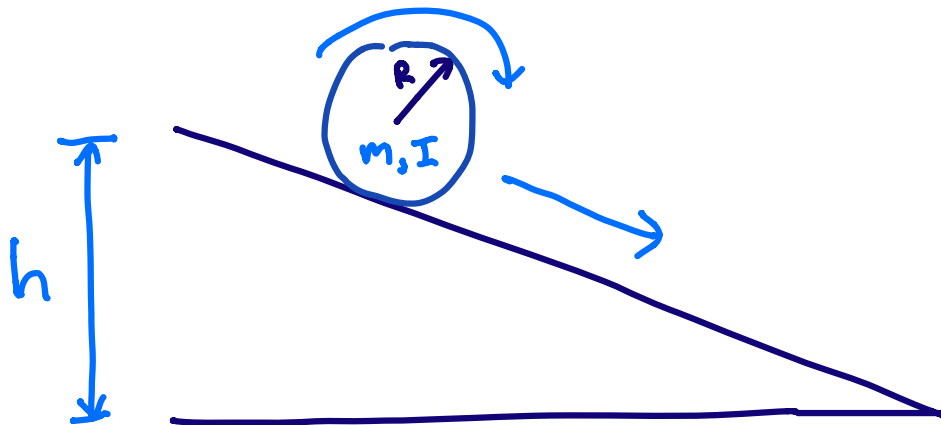


$$I_{\text{stick}} = \int dm \ell^2 = \int_0^L d\ell \underbrace{\frac{m}{L}}_{\text{density}} \ell^2$$
$$= \frac{m}{L} \cdot \frac{L^3}{3} = \frac{1}{3} mL^2$$

$$T_{\text{stick}} = 2\pi \sqrt{\frac{\frac{1}{3} mL^2}{mgL/2}} = 2\pi \sqrt{\frac{2L}{3g}}$$

$$\Rightarrow T_{\text{simp}} > T_{\text{stick}}$$

Rolling without Slipping



Energy Conservation:

$$mgh = \frac{1}{2}mv^2 + \frac{1}{2}I\omega^2$$

initial state only has potential energy

final state only has kinetic energy

final velocity and rotation

define $I = k m R^2$ = numerical constant

$$\Rightarrow mgh = \frac{1}{2}mv^2 + \frac{1}{2}k m R^2 \omega^2 = v^2$$

$$2gh = v^2 (1+k)$$

$$\Rightarrow v = \sqrt{\frac{2gh}{1+k}}$$

<u>type</u>	<u>k</u>	<u>$v/\sqrt{2gh}$</u>
point	0	1
ring	1	$\sqrt{1/2} \approx 0.71$
disc	$\frac{1}{2}$	$\sqrt{2/3} \approx 0.82$
shell	$\frac{2}{3}$	$\sqrt{3/5} \approx 0.77$
ball	$\frac{2}{5}$	$\sqrt{5/7} \approx 0.85$