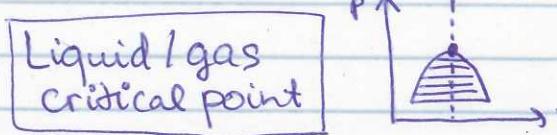


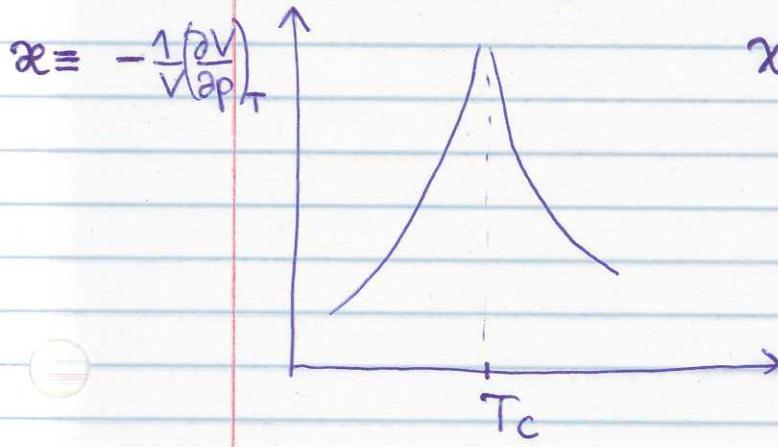
# Macroscopic manifestations of a 2nd-order phase transition and connections with microscopic aspects of a critical behaviour

Recall



Magnetic system

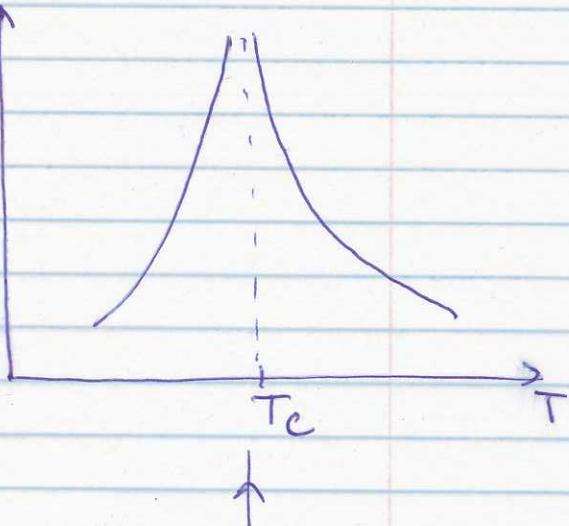
Isothermal compressibility diverges @  $T_c$



tiny change  $\Delta P$   
causes huge  $\Delta V$

$$\chi \equiv \frac{1}{N} \left( \frac{\partial M}{\partial H} \right)_T$$

Spin susceptibility persists



tiny  $\Delta H \Rightarrow$  huge  $\Delta M$

The system is "very soft" & unstable to external perturbations.  
The system "cannot make up its mind" and therefore has large fluctuations

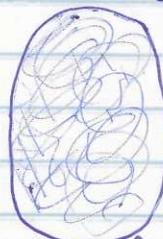
Small fluctuations for  $T > T_c$  and  $T < T_c$ , but large and divergent ("macroscopic") fluctuations as  $T \rightarrow T_c$

$T > T_c$



transparent

$T = T_c$



opaque

critical opalescence

$T < T_c$



transparent

gas

liquid

Fluctuation - dissipation theorem : provides connection between <sup>macroscopic</sup> response functions like susceptibility or compressibility ("dissipation") and microscopic correlations ("fluctuation")

Consider Ising model for concreteness

$$\begin{aligned}
 \chi(T) &= \underset{\text{susceptibility per spin}}{\left( \frac{\partial m}{\partial H} \right)_+} = \underset{\text{arbitrary fixed site}}{\left( \frac{\partial \langle S_0 \rangle}{\partial H} \right)_+} = \\
 &= \frac{\partial}{\partial H} \left( \frac{\sum_{\{S_i\}} S_0 e^{-\beta E_{\text{int}}(S_1, S_2, \dots)} + \beta H \sum_i S_i}{\sum_{\{S_i\}} e^{-\beta E_{\text{int}}(S_1, S_2, \dots)} + \beta H \sum_i S_i} \right) = \\
 &= \langle S_0 \beta \sum_i S_i \rangle - \frac{1}{k_B T} \text{Tr} \{ S_0 e^{-\beta E_{\text{tot}}} \} \text{Tr} \{ \beta \sum_i S_i e^{-\beta E_{\text{tot}}} \} \\
 &= \frac{1}{k_B T} \sum_i \underbrace{(\langle S_0 S_i \rangle - \langle S_0 \rangle \langle S_i \rangle)}_{\Gamma(i)}
 \end{aligned}$$

$\Gamma(i)$  = "connected correlation function"

(Connection with study of "fluctuations" :

$$\boxed{\chi(T) \underset{\text{per spin}}{=} \frac{\langle M^2 \rangle - \langle M \rangle^2}{k_B T \cdot N} \underset{\text{"dissipation"}}{\text{}} \underset{\text{"fluctuation"}}{\text{}}}$$

Thus,

$$\boxed{\chi(T) = \frac{1}{k_B T} \sum_i \Gamma(i) = \frac{1}{k_B T} \int d^d r \Gamma(\vec{r})}$$

arbitrary magnetic system  
on a lattice ; no approximations

continuum  
description

$$\Gamma(\vec{r}) = \langle S_0 S_{\vec{r}} \rangle - \langle S_0 \rangle \langle S_{\vec{r}} \rangle$$

More generally,  $\Gamma(r, r') = \langle S_r S_{r'} \rangle - \langle S_r \rangle \langle S_{r'} \rangle =$

$$= \Gamma(r - r')$$

↑  
for translationally invariant  
systems.

For fluid system

$$\chi(T) = \frac{1}{k_B T} \int d^d \vec{r} G(\vec{r})$$

$\uparrow$   
isothermal  
compressibility

$$G(\vec{r}) = \left[ \langle n(0) n(\vec{r}) \rangle - \langle n(0) \rangle \langle n(\vec{r}) \rangle \right] - \text{density correlation fnctn}$$

$$n(\vec{r}) = \sum_{i=1}^N \delta(\vec{r} - \vec{r}_i)$$

(aka pair correlation fnctn)

Remarks:  $\star$  Expressions for  $\chi, \chi_e$  - examples of "fluctuation-dissipation" theorem, connecting microscopic correlations (fluctuation) with a macroscopic response function (dissipation)

$\star \langle S_0 S_i \rangle$  - can interpret as probability that if spin @ 0 is  $\uparrow$  then spin @ i is  $\uparrow$

$$= (\text{probability both } S_0 \text{ & } S_i \text{ point in the same dir}) - (\text{probability } S_0 \text{ & } S_i \text{ point in opposite dirs.})$$

"measure of correlation"

Is  $S_0$  and  $S_i$  are completely uncorrelated, then

$$\langle S_0 S_i \rangle = \langle S_0 \rangle \langle S_i \rangle, \text{ and } \Gamma(i) = 0$$

In the presence of interactions, spins become correlated. For example, for ferromagnetic interaction spins become positively correlated and  $\Gamma(r) > 0$

\* In order to get  $\chi(\infty) = \infty$ , we must have correlations decaying more slowly than  $\frac{1}{r^{d+\epsilon}}$  at large  $r$ :

otherwise  $\int \Gamma(r) dr < \int \frac{r^{d-1} dr}{r^{d+\epsilon}} = \text{finite}$ .

$\Rightarrow$  at  $T_c$ , correlations are necessarily long-range power law.

Ornstein-Zernike law (correlations in the meanfield theory)

$$T > T_c \quad r \gg \xi \quad \Gamma(r) \sim \frac{e^{-r/\xi}}{r^{\frac{d-1}{2}}}$$

$$\xi(T \neq T_c) \sim |T - T_c|^{-\nu}, \quad \nu_{mf} = \frac{1}{2}$$

$$T = T_c \quad (\text{or } r \ll \xi) \quad \Gamma(r) \sim \frac{1}{r^{d-2}}$$

More generally:

$$\boxed{\Gamma(r) \sim \frac{1}{r^{d-2+\eta}}} \quad \text{for } r \ll \xi$$

$\eta_{mf} = 0$  t modification due to fluctuations

$$\chi(T) \sim \int \frac{r^{d-1} dr}{r^{d-2+\eta}} \sim (\xi(T))^{2-\eta} \sim$$

$$\sim (|T - T_c|)^{-\gamma \cdot (2-\eta)} \sim |T - T_c|^{-\gamma}$$

$$\boxed{\gamma = \gamma \cdot (2-\eta)}$$

meanfield:  $\left. \begin{array}{l} \gamma_{mf} = 1 \\ \gamma_{ms} = \frac{1}{2} \\ \eta = 0 \end{array} \right\} \checkmark$

\* Scattering experiments effectively measure Fourier transform of the density-density correlation function

For example, consider neutron in a medium

$$V(r) = \sum_i w_0 \delta(\vec{r} - \vec{r}_i) = w_0 n(r)$$

↑  
neutron-nucleus interaction; very short-range

Scattering  $\vec{k}_{\text{initial}} \rightarrow \vec{k}_{\text{final}}$  is proportional to

scattering  
from a  
"static  $n(r)$ "  
("snapshot")

$$\left| \langle \vec{k}_f | \hat{V} | \vec{k}_i \rangle \right|^2 = \\ w_0 \int d^3 r e^{i(\vec{k}_i - \vec{k}_f) \cdot \vec{r}} n(\vec{r}) \\ = w_0^2 \int d^3 r d^3 r' n(\vec{r}) n(\vec{r}') e^{+i(\vec{k}_i - \vec{k}_f) \cdot (\vec{r} - \vec{r}')}}$$

Averaging over many neutrons  $\sim$  averaging over possible states of the system  $\sim$  thermal average

$$\begin{aligned} \text{Scattering rate} &\propto \int d^3 r d^3 r' \langle n(r) n(r') \rangle e^{-i \vec{q} \cdot (\vec{r} - \vec{r}')} \\ \text{for } \vec{q} = \vec{k}_f - \vec{k}_i &= \text{Vol.} \int d^3 r G(r) e^{-i \vec{q} \cdot \vec{r}} \end{aligned}$$

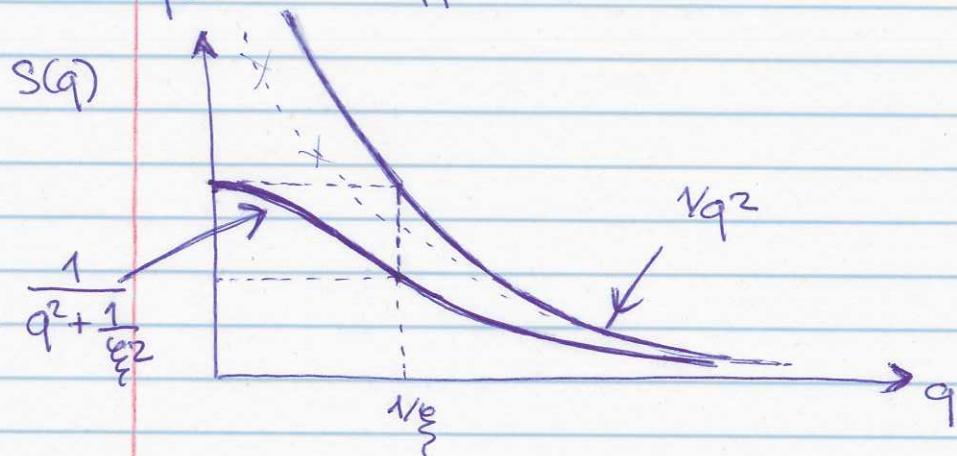
$S(\vec{q})$

$$S(\vec{q}) = \int d^3r G(r) e^{-i\vec{q} \cdot \vec{r}} \quad ] - \text{density-density structure factor in } q\text{-space}$$

Ornstein - Zernike theory:

$$S(\vec{q}) \sim \frac{1}{\vec{q}^2 + \frac{1}{\xi^2}} \quad ] - \text{Lorentzian function of the wavenumber } q.$$

The width of the Lorentzian  $\sim \frac{1}{\xi} \rightarrow 0$  as the critical point is approach



### \* Critical opalescence:

\* "Macroscopic understanding":  $\infty \rightarrow \infty \Rightarrow$  strong fluctuations, regions of various sizes with different densities. Density variations on length scales of the same order as wavelength of the visible light will strongly scatter the light  $\Rightarrow$  the medium looks opaque.

\* "Microscopic understanding": light scattering is determined by  $S(q \sim \frac{2\pi}{\lambda_{\text{light}}})$ . At the critical point

one has fluctuations on all wavelengths, in particular on scales  $\sim \lambda_{\text{light}} \gg$  inter-particle spacing, and  $S(q \sim \frac{2\pi}{\lambda_{\text{light}}})$  is large!