

Pattern Formation in Spatially Extended Systems

Lecture 3: Oscillatory Instabilities

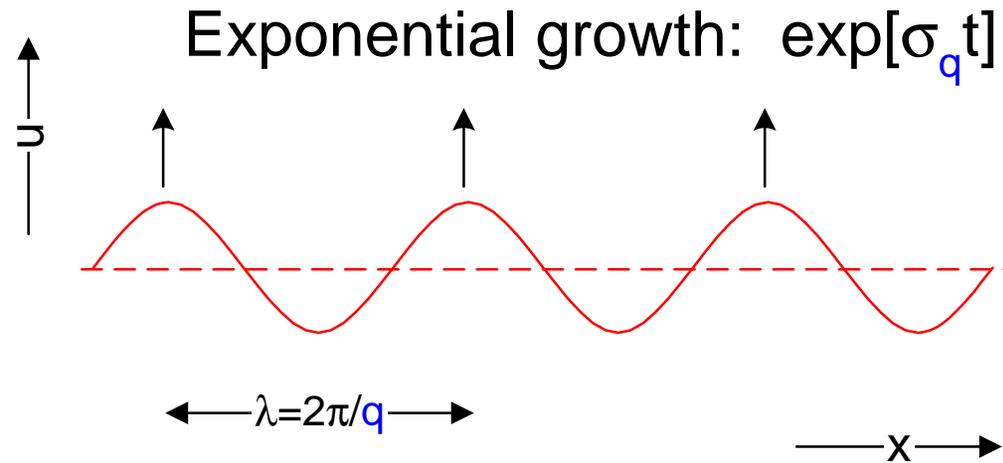
Outline

In this lecture I will discuss pattern formation in oscillatory systems.

As in the stationary case I will introduce the phenomenon using the behavior near an instability from the stationary, spatially uniform state.

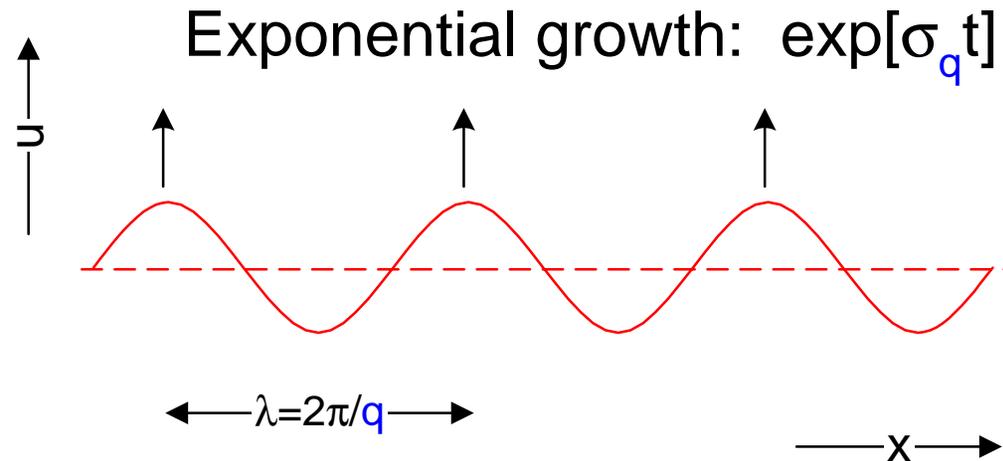
- **Convective v. Absolute Instability**
- **Oscillatory Instability**
 - ◇ CGL equation
 - ◇ Benjamin-Feir instability
 - ◇ Properties of nonlinear waves
 - ◇ Importance of spiral sources in 2d
- **Wave Instability**
 - ◇ Unidirectional wave
 - ◇ Counterpropagating waves

Linear Instability



$$\delta \mathbf{u}_{\mathbf{q}}(\mathbf{x}_{\perp}, z, t) = \mathbf{u}_{\mathbf{q}}(z) e^{i\mathbf{q} \cdot \mathbf{x}_{\perp}} e^{\sigma_{\mathbf{q}} t}$$

Linear Instability

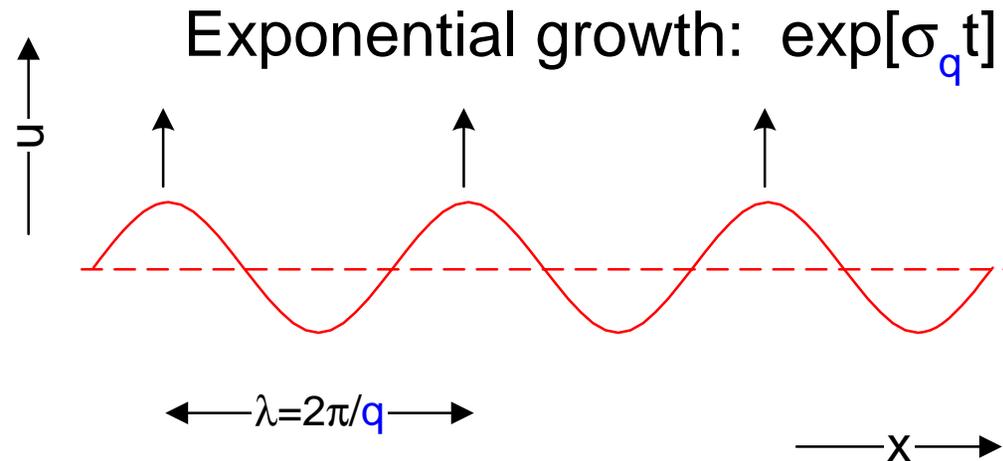


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- for $q_c = 0$: a nonlinear oscillator which also supports travelling waves
- for $q_c \neq 0$: a wave pattern (standing or travelling)

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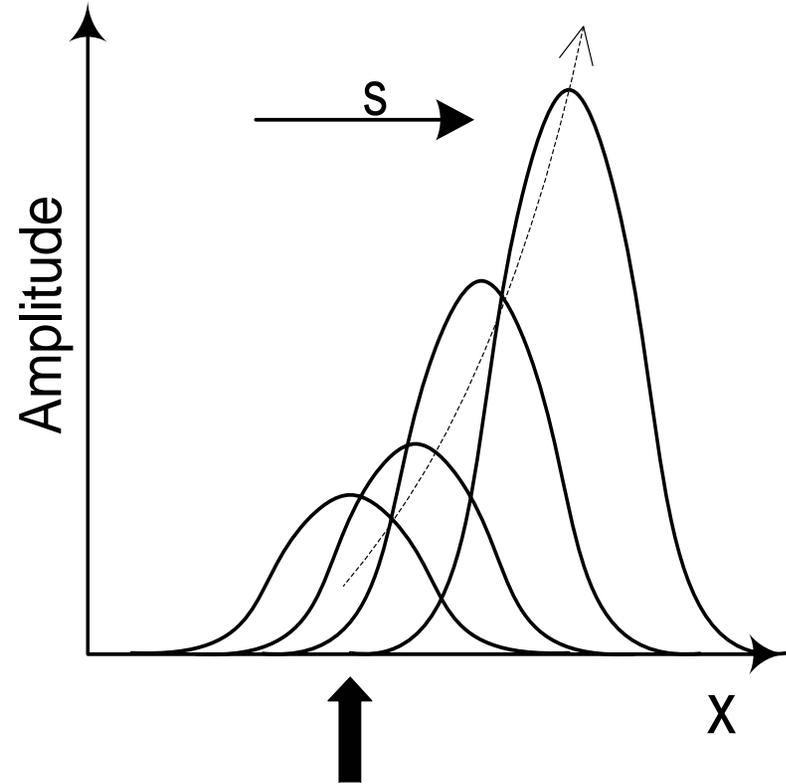
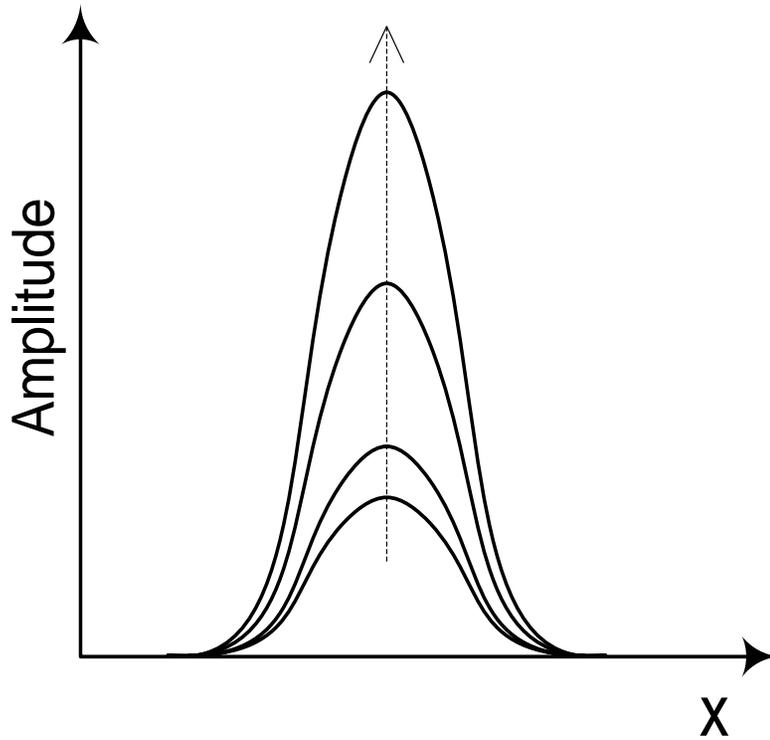
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Important new concept: **absolute** v. **convective instability**

Absolute and convective instability



Conditions for convective and absolute instability

- **Convective instability:** same as condition for instability to Fourier mode
- **Absolute instability:** for a growth rate spectrum σ_q , the system is absolutely unstable if

$$\operatorname{Re} \sigma(\mathbf{q}_s) = 0$$

where \mathbf{q}_s is a *complex* wave vector given by the solution of the stationary phase condition

$$\frac{d\sigma_{\mathbf{q}}}{d\mathbf{q}} = 0$$

Derivation of condition for absolute instability

In the linear regime the disturbance growing from any given initial condition $u_p(\mathbf{x}, t = 0)$ can be expressed as

$$u_p(x, t) = \int_{-\infty}^{\infty} dq e^{iqx + \sigma_q t} \int_{-\infty}^{\infty} dx' u_p(x', 0) e^{-iqx'}$$

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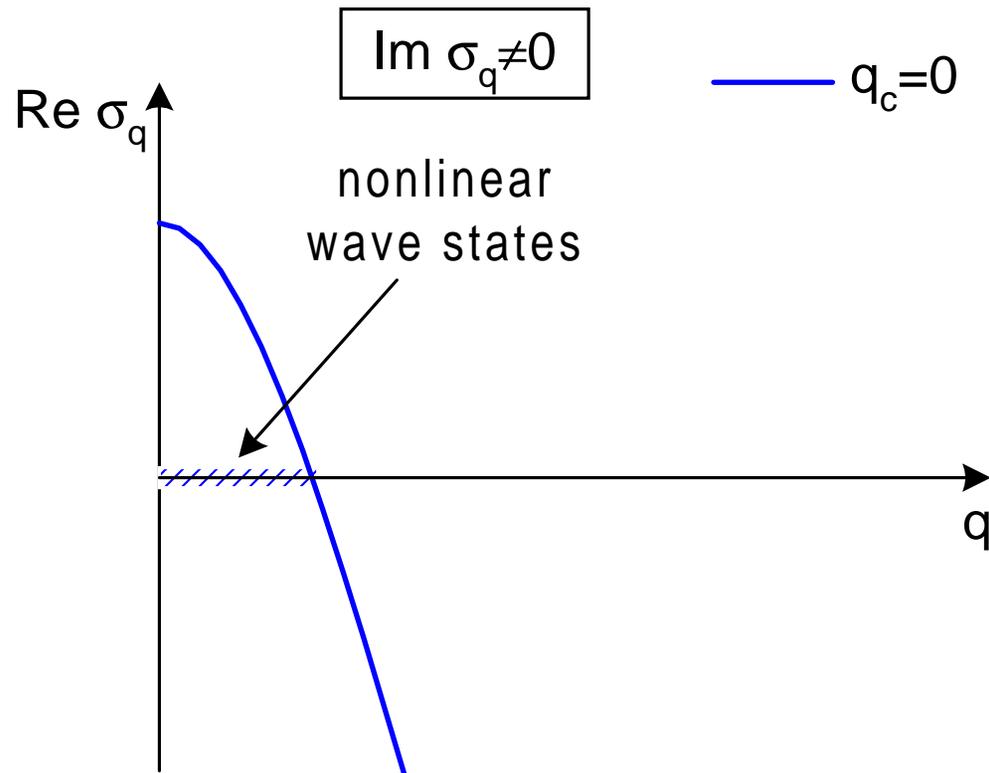
Thus the system will be absolutely unstable for $\text{Re } \sigma_{q_s} > 0$.

Nonlinear oscillators and waves

Insights from amplitude and phase equations

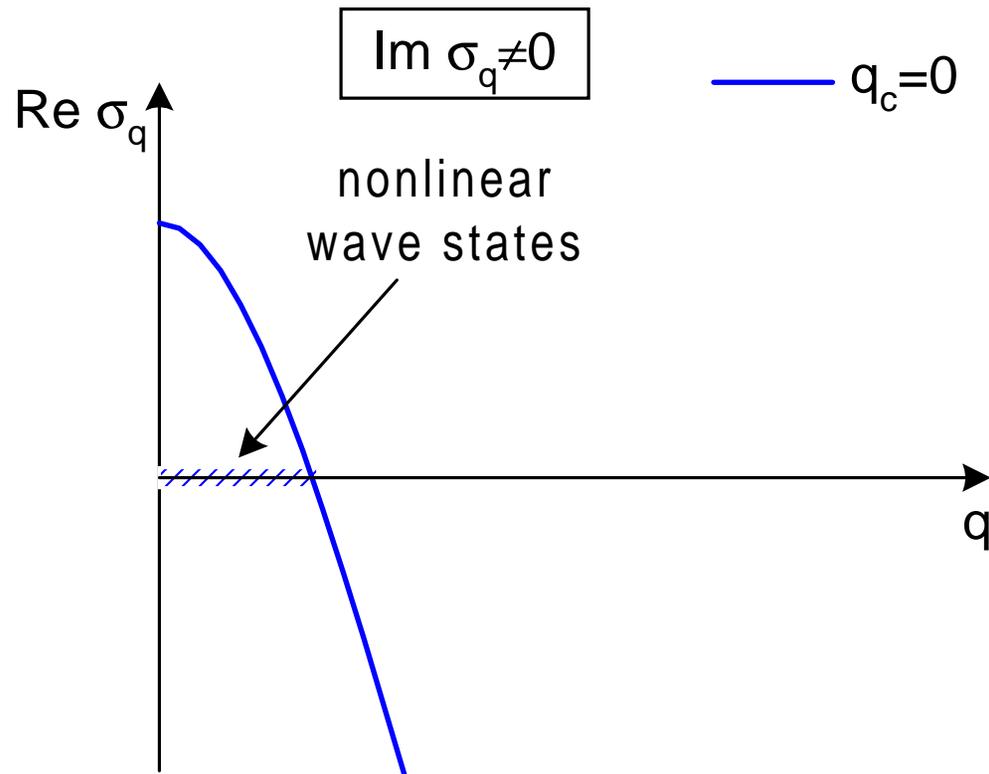
- Oscillatory instability $q_c = 0$
- Wave instability $q_c \neq 0$

Oscillatory instability: Complex Ginzburg-Landau



1d:
$$\partial_T \bar{A} = (1 + ic_0) \bar{A} + (1 + ic_1) \partial_X^2 \bar{A} - (1 - ic_3) |\bar{A}|^2 \bar{A}$$

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$$2d: \quad \partial_T \bar{A} = (1 + ic_0) \bar{A} + (1 + ic_1) \nabla_{\perp}^2 \bar{A} - (1 - ic_3) |\bar{A}|^2 \bar{A}$$

Simulations of the CGL equation

General equation (2d)

$$\partial_T \bar{A} = (1 + ic_0)\bar{A} + (1 + ic_1)\nabla_{\perp}^2 \bar{A} - (1 - ic_3) |\bar{A}|^2 \bar{A}$$

Case simulated: $c_1 = 0$ (choice of parameter), $c_0 = -c_3$ (for simplicity of plots)

$$\partial_T \bar{A} = (1 - ic_3)\bar{A} + \nabla_{\perp}^2 \bar{A} - (1 - ic_3) |\bar{A}|^2 \bar{A}$$

Simulations...

Nonlinear wave patterns

As well as uniform oscillations, CGL equation supports travelling wave solutions, but with properties that are strange to those of us brought up on linear waves:

Nonlinear wave patterns

As well as uniform oscillations, CGL equation supports travelling wave solutions, but with properties that are strange to those of us brought up on linear waves:

- Waves annihilate at shocks rather than superimpose
- Waves disappear at boundaries rather than reflect (not shown)
- Defects: importance as persistent sources
- Spiral defects play a conspicuous role, because they are topologically defined **persistent sources**.
- Instabilities can lead to spatiotemporal chaos

Wave solutions

$$\partial_T \bar{A} = (1 + ic_0)\bar{A} + (1 + ic_1)\nabla_{\perp}^2 \bar{A} - (1 - ic_3) |\bar{A}|^2 \bar{A}$$

Travelling wave solutions

$$\bar{A}_K(\mathbf{X}, T) = a_K e^{i(\mathbf{K} \cdot \mathbf{X} - \Omega_K T)}$$

$$a_K^2 = 1 - K^2 \quad \Omega_K = -(c_0 + c_3) + (c_1 + c_3)K^2$$

Group speed

$$S = d\Omega_K/dK = 2(c_1 + c_3)K$$

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Standing waves, based on the addition of waves at \mathbf{K} and $-\mathbf{K}$ can be constructed, but they are unstable towards travelling waves

Stability analysis

$$\bar{A}_K(\mathbf{X}, T) = (a_K + \delta a)e^{i(\mathbf{K} \cdot \mathbf{X} - \Omega_K T + \delta\theta)}$$

For *small, slowly varying* phase perturbations

$$\partial_T \delta\theta + S \partial_X \delta\theta = D_{\parallel}(K) \partial_X^2 \delta\theta + D_{\perp}(K) \partial_Y^2 \delta\theta$$

with longitudinal and transverse diffusion with constants

$$D_{\parallel}(K) = (1 - c_1 c_3) \frac{1 - \nu K^2}{1 - K^2} \quad D_{\perp}(K) = (1 - c_1 c_3)$$

with

$$\nu = \frac{3 - c_1 c_3 + 2c_3^2}{1 - c_1 c_3}$$

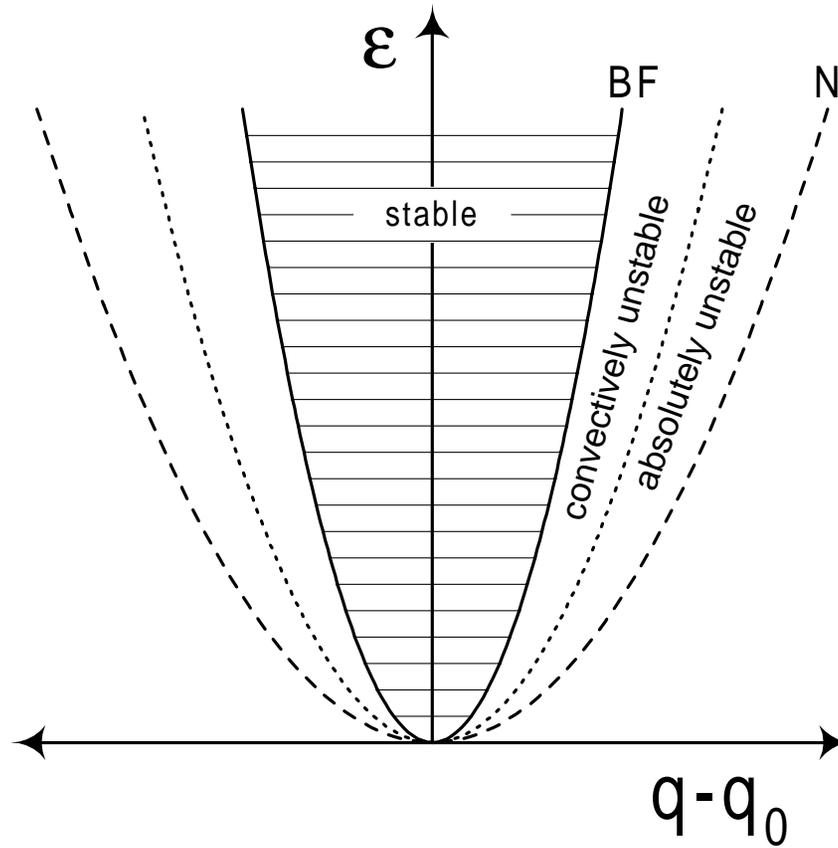
- $D_{\parallel} = 0 \Rightarrow$ Benjamin-Feir instability (longitudinal sideband instability analogous to Eckhaus) for

$$|K| \geq \Lambda_B = \nu^{-1}$$

leaving a stable band of wave numbers with width a fraction ν^{-1} of the existence band.

- For $1 - c_1 c_3 < 0$ *all* wave states are unstable (Newell)

Stability balloon



Shocks: the nonlinear phase equation

For slow phase variations about spatially uniform oscillations (now keeping all terms up to second order in derivatives)

$$\partial_T \theta = \Omega + \alpha \nabla_{\perp}^2 \theta - \beta (\vec{\nabla}_{\perp} \theta)^2$$

with

$$\alpha = 1 - c_1 c_3$$

$$\beta = c_1 + c_3$$

$$\Omega = c_0 + c_3$$

Cole-Hopf transformation

The Cole-Hopf transformation

$$\chi(X, Y, T) = \exp[-\beta\theta(X, Y, T)/\alpha]$$

transforms the nonlinear phase equation into the *linear* equation for χ

$$\partial_T \chi = \alpha \nabla_X^2 \chi$$

Plane wave solutions

$$\chi = \exp [(\mp \beta K X + \beta^2 K^2 T)/\alpha]$$

correspond to the phase variations

$$\theta = \pm K X - \beta K^2 T$$

Cole-Hopf transformation (cont)

Since the χ equation is *linear*, we can superimpose a pair of these solutions

$$\chi = \exp [(-\beta K X + \beta^2 K^2 T)/\alpha] + \exp [(+\beta K X + \beta^2 K^2 T)/\alpha]$$

The phase is

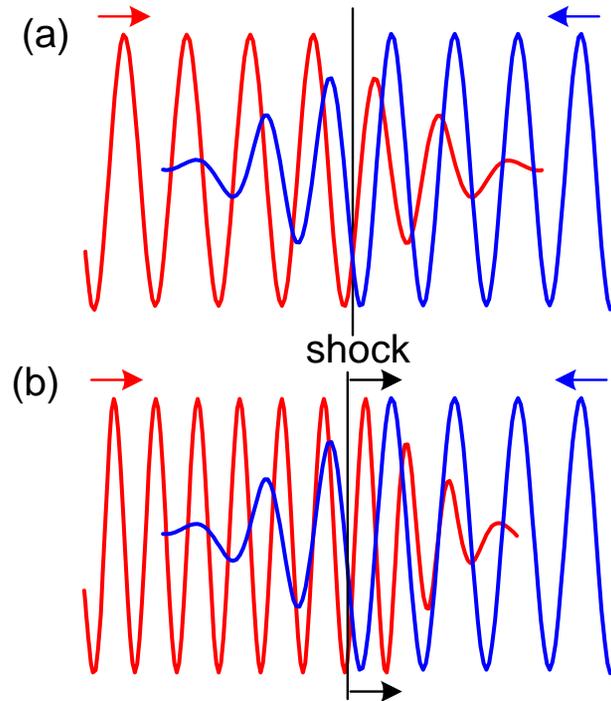
$$\theta = -\beta K^2 T - \frac{\alpha}{\beta} \ln[2 \cosh(\beta K X/\alpha)].$$

For large $|X|$ the phase is given by (assuming βK positive)

$$\theta \rightarrow -K X - \beta K^2 T - \frac{\alpha}{\beta} \exp(-2\beta K X/\alpha) \quad \text{for } X \rightarrow +\infty$$

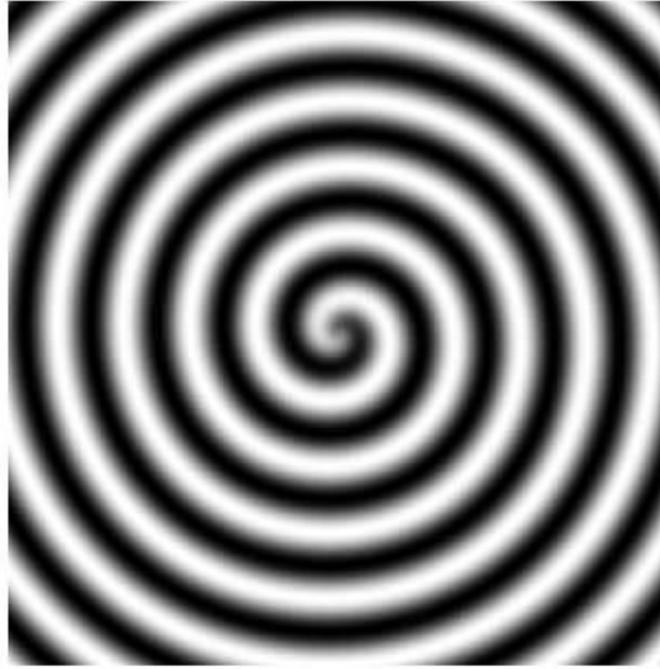
i.e. left moving waves plus exponentially decaying right moving waves with the decay length $\alpha/2\beta K$. Similarly for $X \rightarrow -\infty$ get left moving waves with exponentially small right moving waves.

Shocks



- Shocks are sinks, not sources
- For positive group speed shocks between waves of different frequency move so that the higher frequency region expands

Spiral Defects



m-armed spiral: $\oint \nabla \theta \cdot \mathbf{dl} = m \times 2\pi$

$$\bar{A} = a(R)e^{i(K(R)R+m\theta-\Omega_s T)}$$

with for $R \rightarrow \infty$

$$a(R) \rightarrow a_K \quad K(R) \rightarrow K_s \quad \text{with} \quad \Omega_K(K_s) = \Omega_s$$

Uniqueness

A key question is whether there is a family of spirals giving a continuous range of possible frequencies Ω_s , or there is a unique spiral structure with a prescribed frequency that selects a particular wave number (or possibly a discrete set of possible spirals).

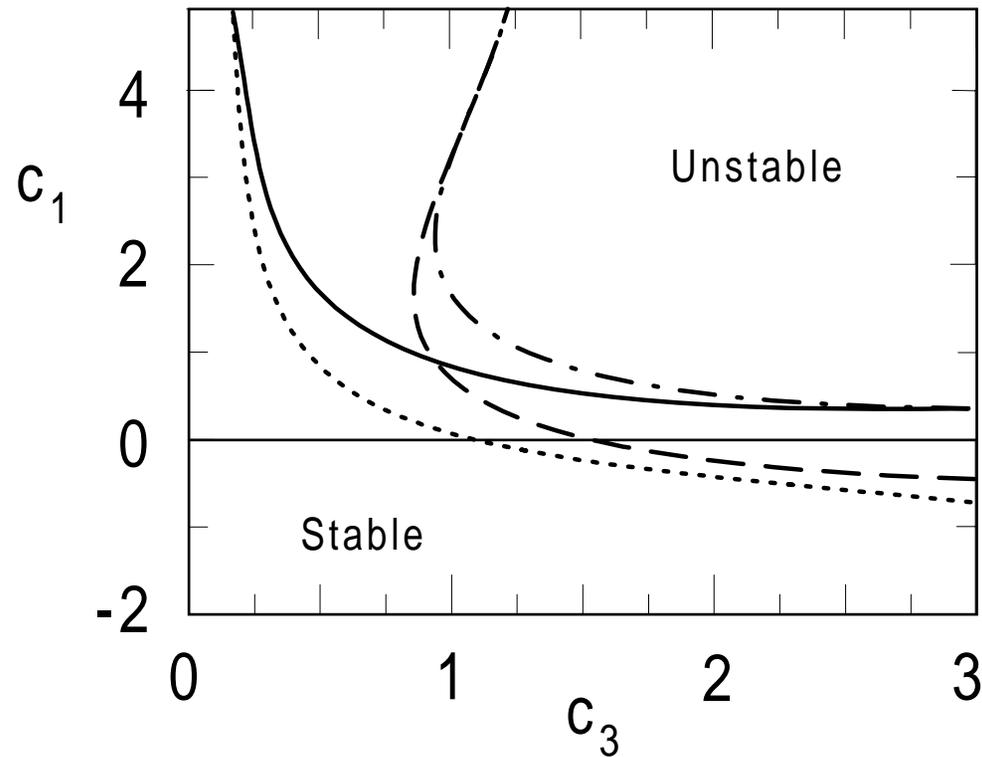
A perturbative treatments of the CGLE for small $c_1 + c_3$ about the real amplitude equation predicts a unique stable spiral structure, with a wave number K_s that varies as (Hagan, 1982)

$$K_s \rightarrow \frac{1.018}{|c_1 + c_3|} \exp\left[-\frac{\pi}{2 |c_1 + c_3|}\right].$$

Stability revisited

- Wave number of nonlinear waves determined by spirals
- Only BF stability of waves at K_s relevant to stability of periodic state
- Convective instability may not lead to breakdown
- Core instabilities may intervene

Stability lines of the CGLE (unstable states are towards larger positive $c_1 c_3$)



solid line: Newell criterion $c_1 c_3 = 1$

dotted line: convective Benjamin-Feir instability of spiral-selected wavenumber

dashed: absolute instability of spiral selected wavenumber

dashed-dotted: absolute instability of whole wavenumber band

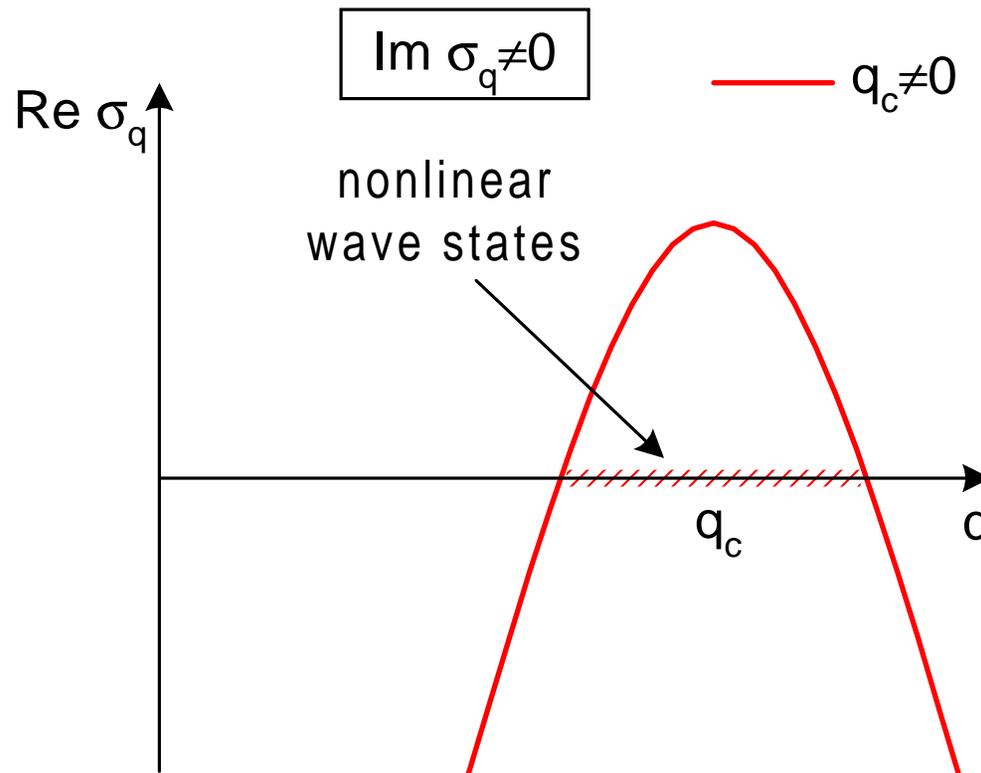
Waves in excitable media

Waves in reaction-diffusion systems such as chemicals or heart tissue show **similar properties**



[From Winfree and Strogatz (1983) and the website of G. Bub, McGill]

Wave instabilities



$$\tau_0(\partial_t + s\partial_x)A = \varepsilon(1 + ic_0)A + \xi_0^2(1 + ic_1)\partial_x^2 A - g_0(1 - ic_3)|A|^2 A$$

No *single* scaling of x , t with ε eliminates the small parameter ε from equation.

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- Lowest order equation is just the propagation

$$\partial_{T_p} \bar{A} + s \partial_X \bar{A} = 0.$$

Solution is that \bar{A} is a function of the reduced coordinate $\xi = X - sT_p$, i.e.

$\bar{A}(X, T_p, T) = \bar{A}(\xi, T)$. Physically this corresponds to transforming to a frame moving at the group speed s .

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- Then use

$$\partial_\xi \rightarrow \varepsilon^{-1/2} \partial_x, \quad \partial_T \rightarrow \varepsilon^{-1} (\partial_t + s \partial_x).$$

Solutions

$$(\partial_t + s\partial_x)A = \varepsilon(1 + ic_0)A + \xi_0^2(1 + ic_1)\partial_x^2 A - g_0(1 - ic_3)|A|^2 A$$

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Also can treat counterpropagating waves for case with $x \rightarrow -x$ symmetry

$$(\partial_T + S\partial_X)\bar{A}_R = (1 + ic_0)\bar{A}_R + (1 + ic_1)\partial_X^2 \bar{A}_R - (1 - ic_3)|\bar{A}_R|^2 \bar{A}_R - g_1(1 - ic_2)|\bar{A}_L|^2 \bar{A}_R$$

$$(\partial_T - S\partial_X)\bar{A}_L = (1 + ic_0)\bar{A}_L + (1 + ic_1)\partial_X^2 \bar{A}_L - (1 - ic_3)|\bar{A}_L|^2 \bar{A}_L - g_1(1 - ic_2)|\bar{A}_R|^2 \bar{A}_L$$

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- (Knobloch and de Luca 1990) For $s = O(1)$ and ε small, interaction with inhomogeneity in medium or counterpropagating wave is *nonlocal* e.g.

$$|\bar{A}_L|^2 \bar{A}_R \rightarrow \left(L^{-1} \int |\bar{A}_L|^2 dX \right) \bar{A}_R$$

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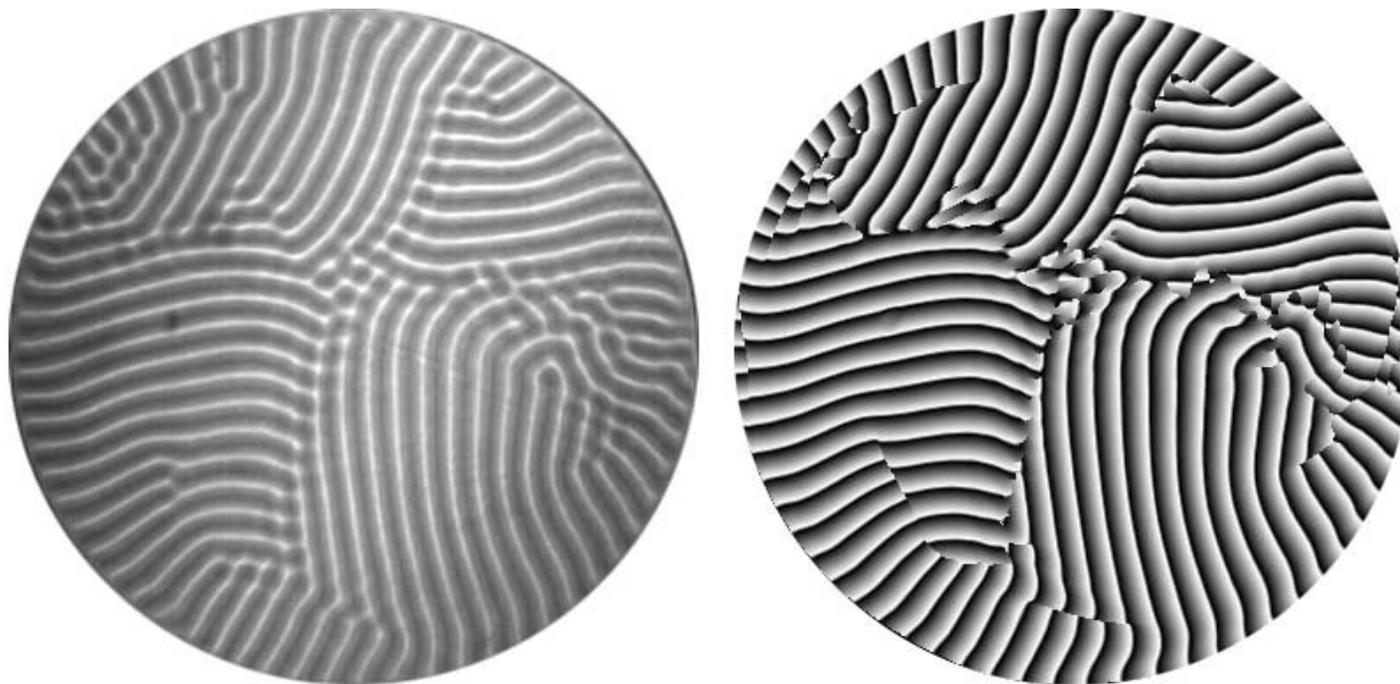
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- (Martel and Vega 1998) “Hyperbolic” equations without ∂_X^2 terms

Wave instability in 2d



[From La Porta and Surko (1998)]

Other issues for wave instabilities

- **Noise sustained structures** in convectively unstable domain
- Global modes (e.g. for $\varepsilon(x)$): local absolutely unstable region sustains disturbance in convectively unstable region (Chomaz et al., 1988)
- Complex dynamics of counterpropagating waves in finite geometry (e.g. **blinking states**)

Conclusions

In this lecture I discussed pattern formation in oscillatory systems. Some of the key concepts were:

- **Convective v. Absolute Instability**
- **Oscillatory Instability**
 - ◇ CGL equation
 - ◇ Benjamin-Feir instability
 - ◇ Properties of nonlinear waves
 - ◇ Importance of spiral sources in 2d
- **Wave Instability**
 - ◇ Importance of propagation term