

1. Time-dependent Perturbation Theory.

Find the time dependent dipole moment using:

$$\mu(t) = \langle \Psi(r, t) | ez | \Psi(r, t) \rangle$$

Substitute in the expression for the time-dependent wavefunction:

$$\Psi(r, t) = a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} + \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar}$$

This substitution yields:

$$\mu(t) = \left\langle a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} + \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \left| ez \right| a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} + \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \right\rangle$$

We expand the integral into four terms:

$$\begin{aligned} \mu(t) = & \left\langle a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} \left| ez \right| a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} \right\rangle + \left\langle a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} \left| ez \right| \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \right\rangle \\ & + \left\langle \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \left| ez \right| a_0(t)\varphi_0^{(0)}(r)e^{-iE_0t/\hbar} \right\rangle + \left\langle \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \left| ez \right| \sum_{n \neq 0}^{\infty} a_n(t)\varphi_n^{(0)}(r)e^{-iE_n t/\hbar} \right\rangle \end{aligned}$$

Substitute in the following expression for a_n :

$$a_n(t) = \frac{-1}{\hbar} \langle \varphi_n^{(0)} | H^{(1)} | \varphi_0^{(0)} \rangle \left[\frac{e^{i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right]$$

Note that we use the time independent perturbation $H^{(1)}$ from:

$$H^{(1)}(t) = H^{(1)} 2 \cos(\omega t) = e\hat{z}E 2 \cos(\omega t)$$

This results in a new expression for the dipole moment:

$$\begin{aligned}
\mu(t) &= |a_0(t)|^2 \langle \varphi_0^{(0)}(r) | e_z | \varphi_0^{(0)}(r) \rangle \\
&+ \frac{-a_0(t)^*}{\hbar} \sum_{n \neq 0}^{\infty} e^{i(E_0 - E_n)t/\hbar} \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \left[\frac{e^{i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \\
&+ \frac{-a_0(t)}{\hbar} \sum_{n \neq 0}^{\infty} e^{-i(E_0 - E_n)t/\hbar} \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \langle \varphi_n^{(0)}(r) | e_z | \varphi_0^{(0)}(r) \rangle \left[\frac{e^{-i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{-i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \\
&+ \left\langle \sum_{n \neq 0}^{\infty} \left[\frac{e^{-i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{-i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \varphi_n^{(0)}(r) \frac{e^{-iE_n t/\hbar}}{\hbar} \right| \\
&e_z \left| \sum_{n \neq 0}^{\infty} \left[\frac{e^{-i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{-i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \varphi_n^{(0)}(r) \frac{e^{-iE_n t/\hbar}}{\hbar} \right\rangle
\end{aligned}$$

By ignoring terms that are not linear in E_z (because only terms linear in $E_0 = E_z$ can be linear in $E(t) = 2E_0 \cos \omega t$) and noting that $\omega_{no} \hbar = E_n - E_0$ we obtain:

$$\begin{aligned}
\mu(t) &= \frac{-1}{\hbar} \sum_{n \neq 0}^{\infty} e^{i\omega_{no}t} \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \left[\frac{e^{i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \\
&+ \frac{-1}{\hbar} \sum_{n \neq 0}^{\infty} e^{-i\omega_{no}t} \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \langle \varphi_n^{(0)}(r) | e_z | \varphi_0^{(0)}(r) \rangle \left[\frac{e^{-i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{-i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \\
&= 2\text{Re} \left[\frac{-1}{\hbar} \sum_{n \neq 0}^{\infty} e^{i\omega_{no}t} \langle \varphi_n^{(0)}(r) | E_z e_z | \varphi_0^{(0)}(r) \rangle \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \left[\frac{e^{i(\omega_{no} + \omega)t} - 1}{\omega_{no} + \omega} + \frac{e^{i(\omega_{no} - \omega)t} - 1}{\omega_{no} - \omega} \right] \right] \\
&= \frac{-2E_z}{\hbar} \text{Re} \left[\sum_{n \neq 0}^{\infty} \left| \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \right|^2 \left[\frac{\omega_{no} \left[(e^{i\omega t} + e^{-i\omega t}) - (2e^{-i\omega_{no}t}) \right] - \omega (e^{i\omega t} - e^{-i\omega t})}{\omega_{no}^2 - \omega^2} \right] \right] \\
&= \frac{-2E_z}{\hbar} \text{Re} \left[\sum_{n \neq 0}^{\infty} \frac{\left| \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \right|^2}{\omega_{no}^2 - \omega^2} [2\omega_{no} \cos \omega t - 2\omega_{no} (\cos \omega_{no} t + i \sin \omega_{no} t) - i2\omega \sin \omega t] \right] \\
&= \frac{-2E_z}{\hbar} \sum_{n \neq 0}^{\infty} \frac{\left| \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \right|^2}{\omega_{no}^2 - \omega^2} [2\omega_{no} \cos \omega t - 2\omega_{no} \cos \omega_{no} t]
\end{aligned}$$

Since we're only interested in the term which is linear in $E(t) = 2E_0 \cos \omega t$, we ignore the second term and get:

$$\begin{aligned}
\mu(t) &= \frac{-2E_z}{\hbar} \sum_{n \neq 0}^{\infty} \frac{\left| \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \right|^2}{\omega_{no}^2 - \omega^2} 2\omega_{no} \cos \omega t \\
&= \frac{-2E(t)}{\hbar} \sum_{n \neq 0}^{\infty} \frac{\omega_{no} \left| \langle \varphi_0^{(0)}(r) | e_z | \varphi_n^{(0)}(r) \rangle \right|^2}{\omega_{no}^2 - \omega^2}
\end{aligned}$$

Finally, we find $\alpha(\omega)$ using:

$$\alpha(\omega) = \frac{\mu}{E(t)}$$

Substituting our expression for the dipole moment in terms of $E(t)$ gives our final result:

$$\alpha(\omega) = \frac{-2}{\hbar} \sum_{n \neq 0}^{\infty} \frac{-2\omega_{n0}}{\hbar} \frac{|\langle \varphi_0^{(0)}(r) | ez | \varphi_n^{(0)}(r) \rangle|^2}{\omega_{n0}^2 - \omega^2}$$

2. Hydrogen in a Magnetic Field.

- a. The motion of an electron about the nucleus constitutes an electric current, which produces a dipole moment that can interact with an external magnetic field. For the motion of the electron, we know $vt = 2r\pi$, where v is the speed of the electron, t is the time it takes to make one revolution, and r is the radius of the orbit.

For a dipole moment, $\mu = IA$. Since $I = \frac{q}{t}$, we can write

$$\mu = \frac{qv}{2\pi r} \pi r^2 = \frac{qvr}{2}.$$

Writing μ as a vector,

$$\vec{\mu} = \frac{q(\vec{r} \times \vec{v})}{2} = \frac{q(\vec{r} \times \vec{p})}{2m} = \frac{q\vec{L}}{2m}.$$

We know:

$$|L| = \hbar[l(l+1)]^{\frac{1}{2}},$$

So:

$$|\mu| = \frac{|e|\hbar}{2m} [l(l+1)]^{\frac{1}{2}} = \beta_e [l(l+1)]^{\frac{1}{2}},$$

where β_e is the Bohr magneton:

$$\beta_e = \frac{|e|\hbar}{2m_e}.$$

Its units are $\frac{J}{T}$.

- b. The potential energy of a dipole in a magnetic field is $U = -\mu \cdot B$. Since our magnetic field is pointing in the z direction, only the z component of μ will interact with it. The new Hamiltonian will just be the one for hydrogen with no magnetic field plus the new potential due to this interaction:

$$\hat{H} = \hat{H}_0 + U = \hat{H}_0 - B_z \mu_z = \hat{H}_0 + \frac{|e|\hbar}{2m} \hat{L}_z B_z = \hat{H}_0 + \frac{\beta_e B_z}{\hbar} \hat{L}_z$$

Because the hydrogenlike wavefunctions are eigenfunctions of \hat{H}_0 and \hat{L}_z , they will still be eigenfunctions for this Hamiltonian. Applying the Hamiltonian to the wavefunctions gives:

$$\hat{H}\psi_{nlm} = \hat{H}_0\psi_{nlm} + \frac{\beta_e B_z}{\hbar} \hat{L}_z \psi_{nlm},$$

So

$$E\psi_{nlm} = E_n^0\psi_{nlm} + \beta_e B_z m \psi_{nlm}$$

And:

$$E = E_n^0 + \beta_e B_z m.$$

- c. Because $m = 0$ for s orbitals, the energies of the $1s$ and $2s$ will be unchanged by the magnetic field. The $2p$, however, will become degenerate. The energy of the $m = 1$ orbital will be raised by $\beta_e B_z$, while the $m = -1$ will be lowered by the same factor.

For a magnetic field of magnitude 15T, the splitting will be 1.391×10^{-22} J. The difference in energy between unperturbed $1s$ and $2p$ hydrogen orbitals is 1.635×10^{-18} J, so the splitting is about .01% of the energy difference.

- d. The $l = 2$ state of hydrogen has 5 possible values for m , while the $l = 3$ state has 7. So the total number of possible transitions is 35.

If $\Delta m = 0$, each orbital has only one choice of where it can go, so the possible number of transitions is 5. If $\Delta m = \pm 1$, each orbital has two possible destinations, so there are 10 possible transitions.

The quantum number m indicates the projection of angular momentum along the z axis, and therefore the component of motion of the electron in the xy plane. If the electric field vector of the light is parallel to the external magnetic field, and, therefore, L_z , it is perpendicular to the component of motion in the xy plane and will not be able to exert any torque to change it. So the value of the z projection of angular momentum, and hence m , will not be changed.

If the electric field vector of the light is perpendicular to the z axis, it can effect the motion of the electron in the xy plane and therefore change L_z and m .